

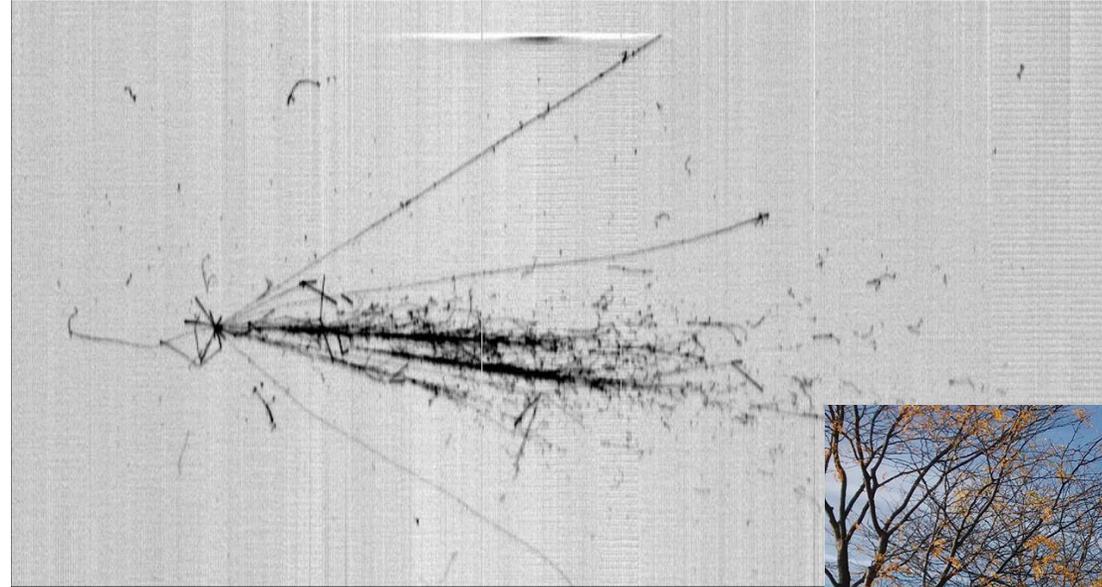
# *ICARUS: a new challenge within the Fermilab Short Baseline Neutrino Program*

*Andrea Rappoldi  
INFN Pavia*

*on behalf of the  
**ICARUS** collaboration*

*New Trends in High-  
Energy Physics*

*Odessa  
May 12-18, 2019*



- ICARUS-T600 at LNGS: *a very successful run*
- Sterile neutrinos and the SBN project
- ICARUS upgrade: *from underground up to the surface*
- Installation and commissioning at **FNAL**

*O(100) scientists*



Italy

*Catania (INFN and Univ.)  
GSSI  
LNGS  
INFN Milano Bicocca  
INFN Napoli  
Padova (INFN and Univ.)  
Pavia (INFN and Univ.)*

*Spokesman: C. Rubbia (GSSI)*



U.S.A.

*Brookhaven (BNL)  
Colorado State  
FNAL  
Houston  
Pittsburgh  
Rochester  
SLAC  
Texas (Arlington)  
Southern Methodist Univ.*



Mexico

*CINVESTAV*



Neutrino Platform NP01



*INTENSE*

This work was supported by the EU *Horizon 2020* Research and Innovation Programme under the Marie Skłodowska-Curie Grant Agreement No. 822185

# Liquid Argon TPC: an “electronic bubble chamber”

4

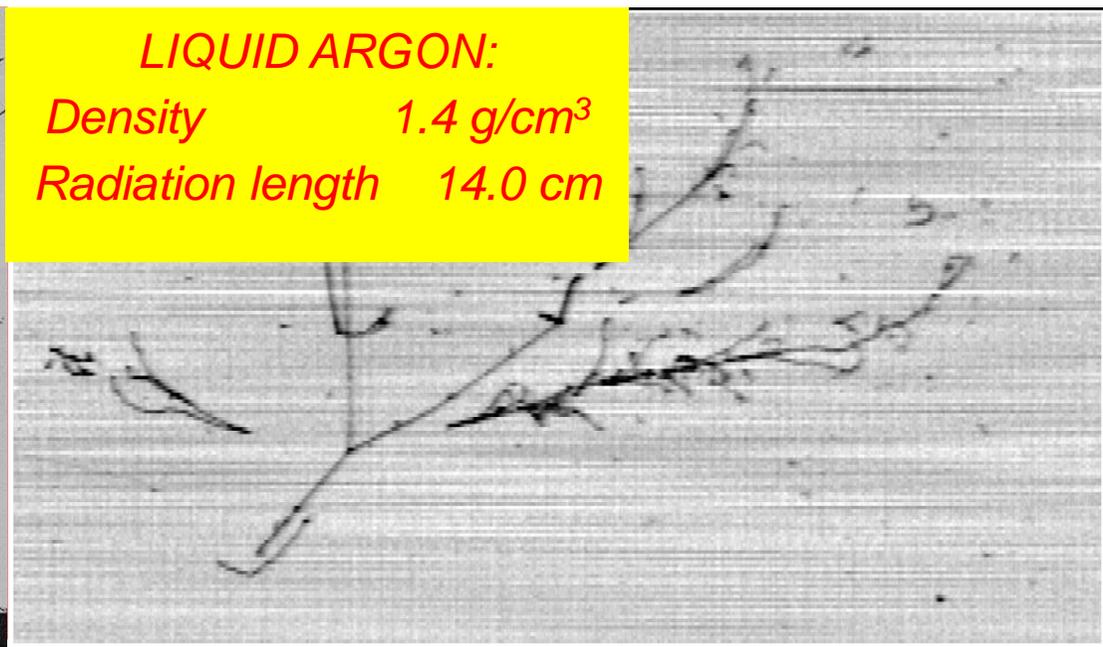
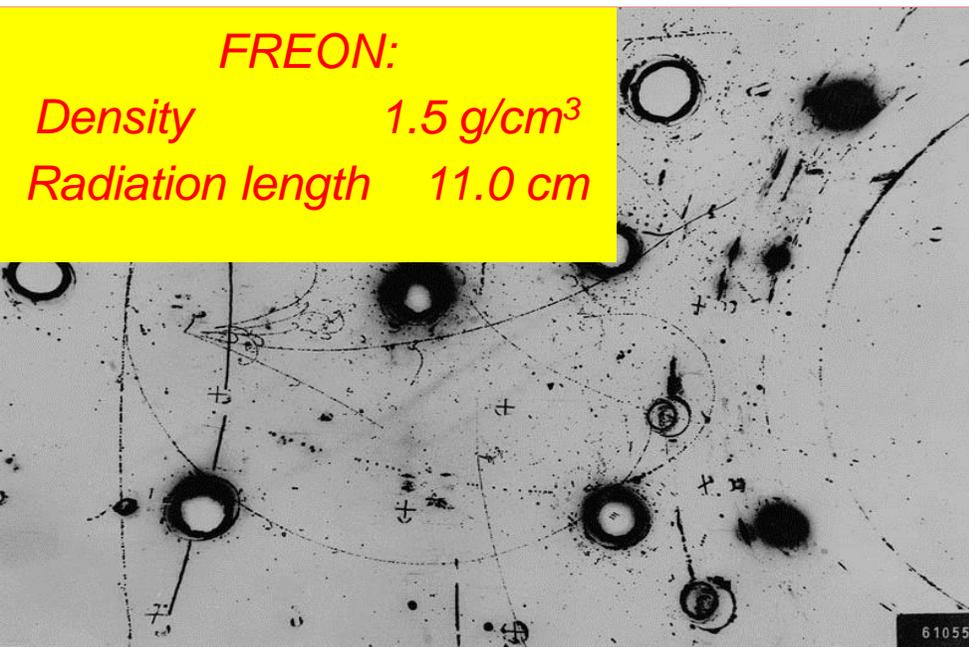
- First proposed by C. Rubbia in 1977: long R&D at INFN and CERN culminated in first large-scale experiment: **ICARUS-T600** at LNGS (2010)
- LAr TPCs are suitable detectors for neutrino physics:
  - 3D reconstruction with high ( $\text{mm}^3$ ) spatial granularity
  - Homogeneous, full-sampling calorimetry for contained particles
  - Scintillation light  $\rightarrow$  fast signals for timing/triggering
  - Electrons can drift for several meters  $\rightarrow$  very large detectors (ktons)
- LAr physical parameters very similar to Freon of “classic” bubble chambers:

## *FREON:*

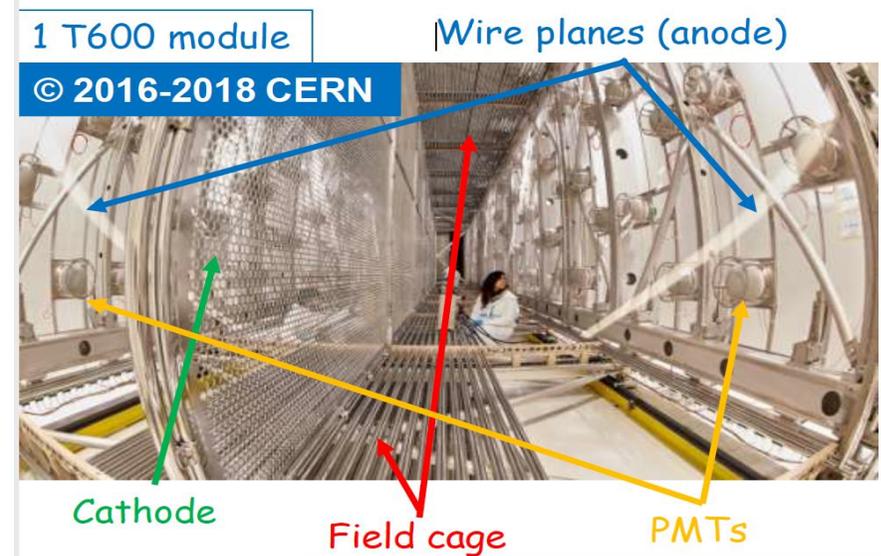
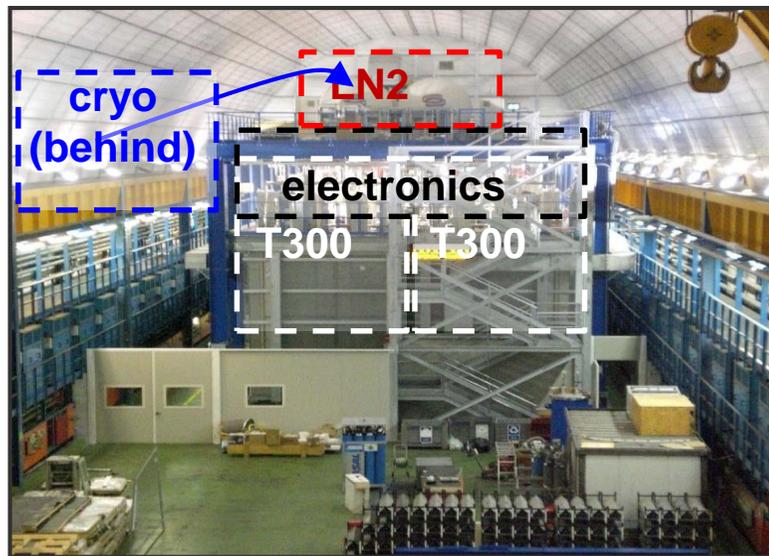
Density  $1.5 \text{ g/cm}^3$   
Radiation length  $11.0 \text{ cm}$

## *LIQUID ARGON:*

Density  $1.4 \text{ g/cm}^3$   
Radiation length  $14.0 \text{ cm}$



- 2 identical modules: each is  $19.6 \times 3.6 \times 3.9 \text{ m}^3$ , with total active mass of **760 t**
- Drift distance 1.5 m and electric field 500 V/cm  $\rightarrow$  drift time  $\sim 1 \text{ ms}$
- 3 signal wire planes with "non-destructive" readout for a total of  $\sim$  **54000 channels**
- Pitch and inter-plane distance both 3 mm; 400 ns sampling time
- 74 (20+54) 8" **PMTs** with TPB wavelength-shifter coating



- ICARUS was exposed to CNGS beam and cosmics for 3 years (2010-2013)
- Run confirmed expected performance and obtained important physics results
- It proved the maturity of the LAr-TPC technique for large-scale experiments

***ICARUS paved the way to the next generation long-baseline project: DUNE***

- High electron lifetime:  $> 7$  ms (impurity concentration  $< 40$  ppt)  
Crucial step towards future larger detectors

2014 JINST 9 P12006

- Excellent spatial/calorimetric reconstruction.

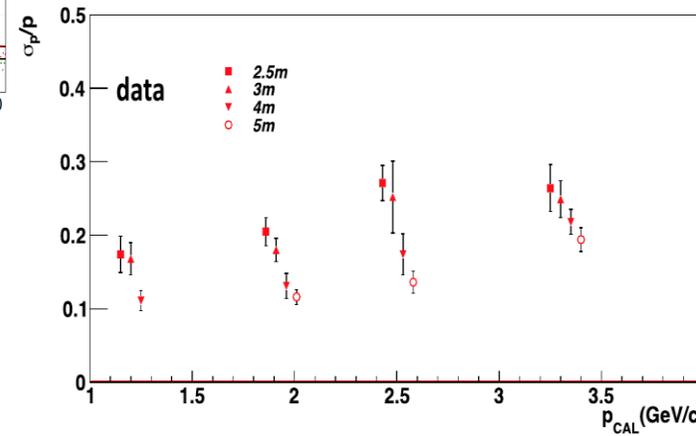
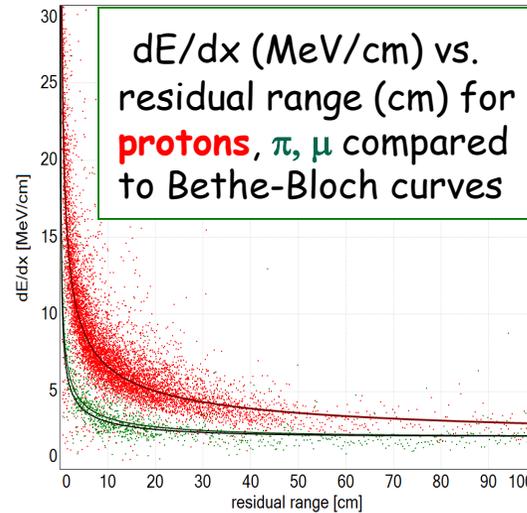
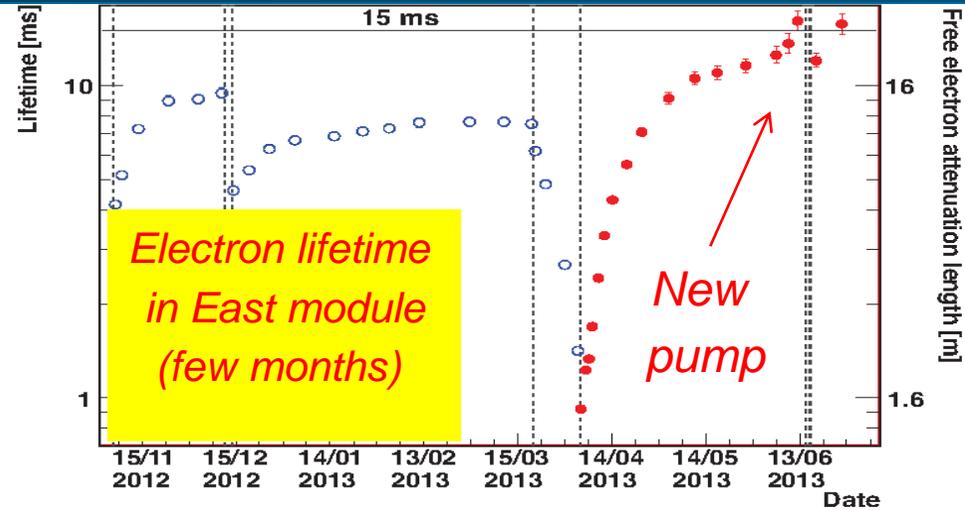
Accurate  $dE/dx$  measurement with fine sampling ( $0.02X_0$ ).

Particle ID from  $dE/dx$  vs. range

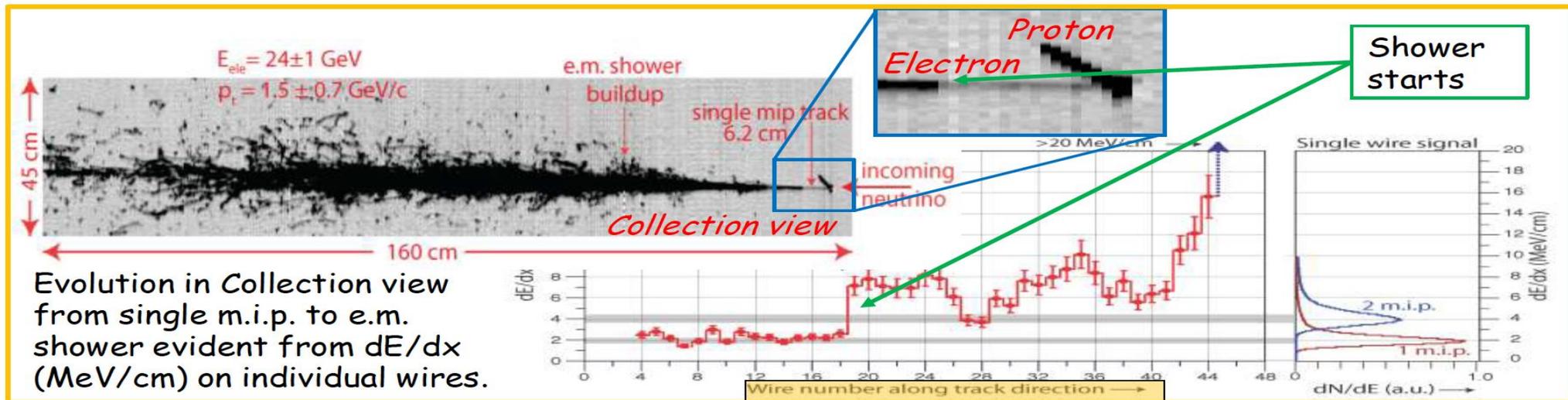
AHEP (2013) 260820

- Momentum of escaping muons can be measured by multiple Coulomb scattering. Average  $\sim 15\%$  resolution on stopping muons ( $0.5 \div 5$  GeV/c)

JINST 12P04010

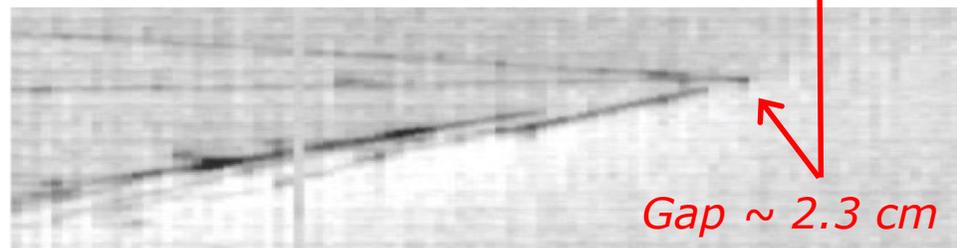
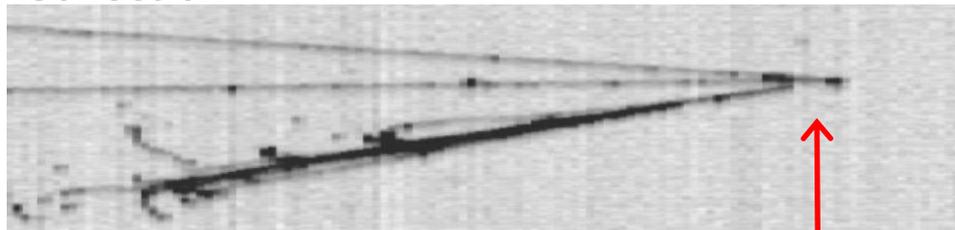


## High-energy CNGS $\nu_e$ CC interaction:



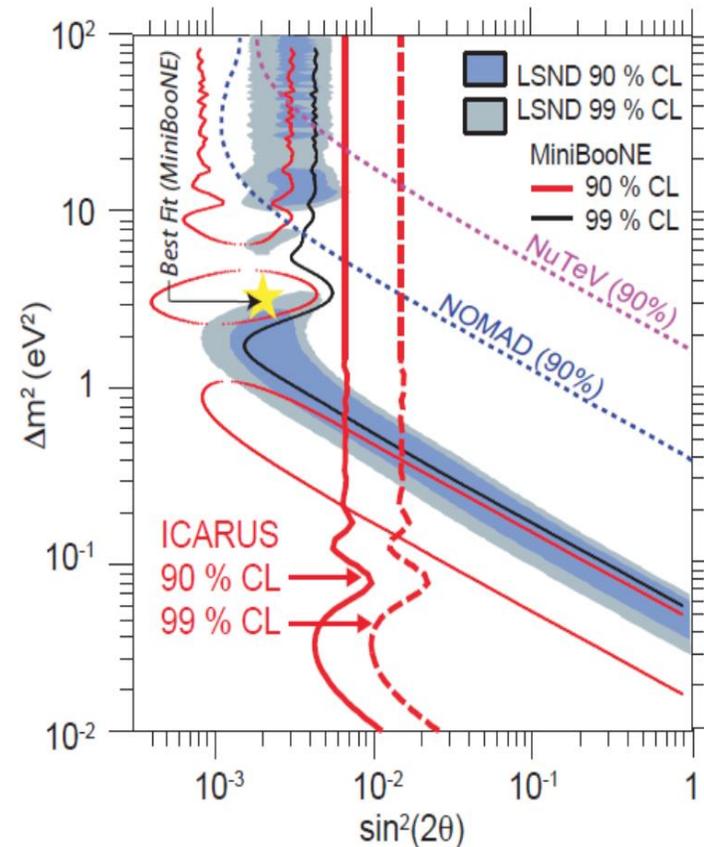
- $\nu_e$  CC event (electron-initiated EM showers) separation from NC background with  $\pi^0$  ( $\gamma$ -initiated showers) is crucial for oscillation physics
- LAr-TPC provides 3 handles:
  - Visual identification of  $\gamma$  conversion gap
  - Reconstruction of  $\pi^0$  invariant mass
  - $dE/dx$  measuring on each wire allows calorimetric accuracy with fine sampling (2%  $X_0$ )

Collection



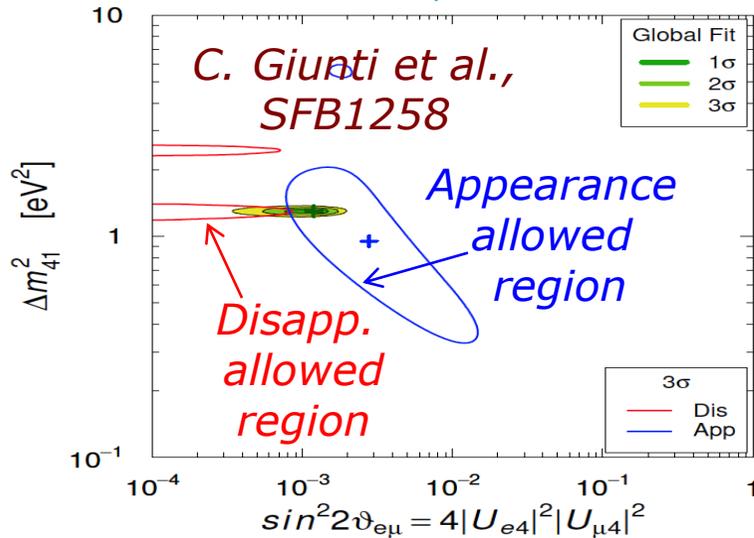
Induction 2

- ICARUS searched for sterile  $\nu$  oscillations through  $\nu_e$  appearance in the CNGS beam
- $L/E \sim 36 \text{ km/GeV}$ , far from LSND value  $\sim 1 \text{ km/GeV}$   
 $\rightarrow$  "sterile-like" oscillation was averaged out, canceling energy dependence
- $7.9 \cdot 10^{19}$  pots analyzed ( $\sim 2650$   $\nu$  interactions)
- Expected  $\sim 8.5 \pm 1.1$   $\nu_e$  background events in absence of anomaly, mostly from intrinsic  $\nu_e$  beam contamination
- Estimated  $\nu_e$  identification efficiency  $\sim 74\%$  with negligible background from misidentification
- **7 events observed  $\rightarrow$  no evidence of oscillation**
- Most of LSND allowed region is excluded - except for small area around  $\sin^2 2\theta \sim 0.005$ ,  $\Delta m^2 < 1 \text{ eV}^2$
- Similar result by OPERA with same CNGS beam and different detection technique

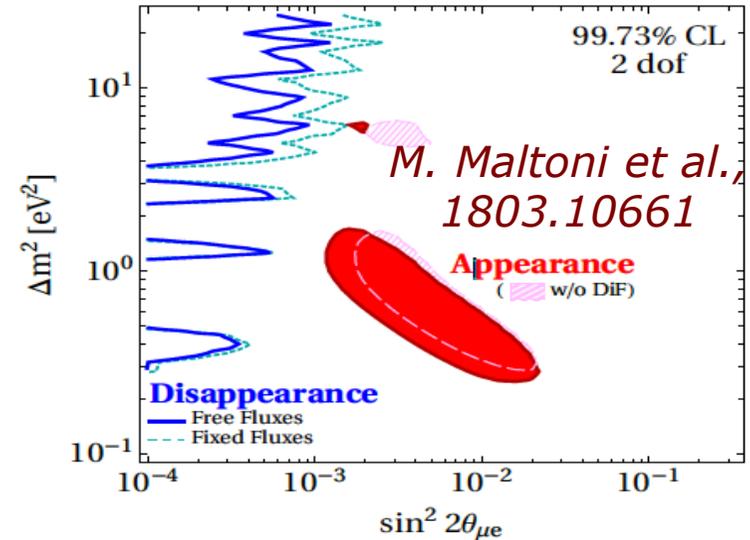


*Eur. Phys. J. C*  
 (2013) 73:2599

- The sterile neutrino scenario is far from understood and needs a definitive clarification
- Some “anomalies” from accelerators (LSND), reactor, neutrino sources, point out to flavour transitions in the  $\Delta m^2 \sim 1 \text{ eV}^2$  range
- However, no evidence of oscillations in  $\nu_\mu$  disappearance data (MINOS, IceCube)
- Tension between appearance and  $\nu_\mu$  disappearance results. **Measuring both channels with the same experiment will help disentangle**



Combined analyzes



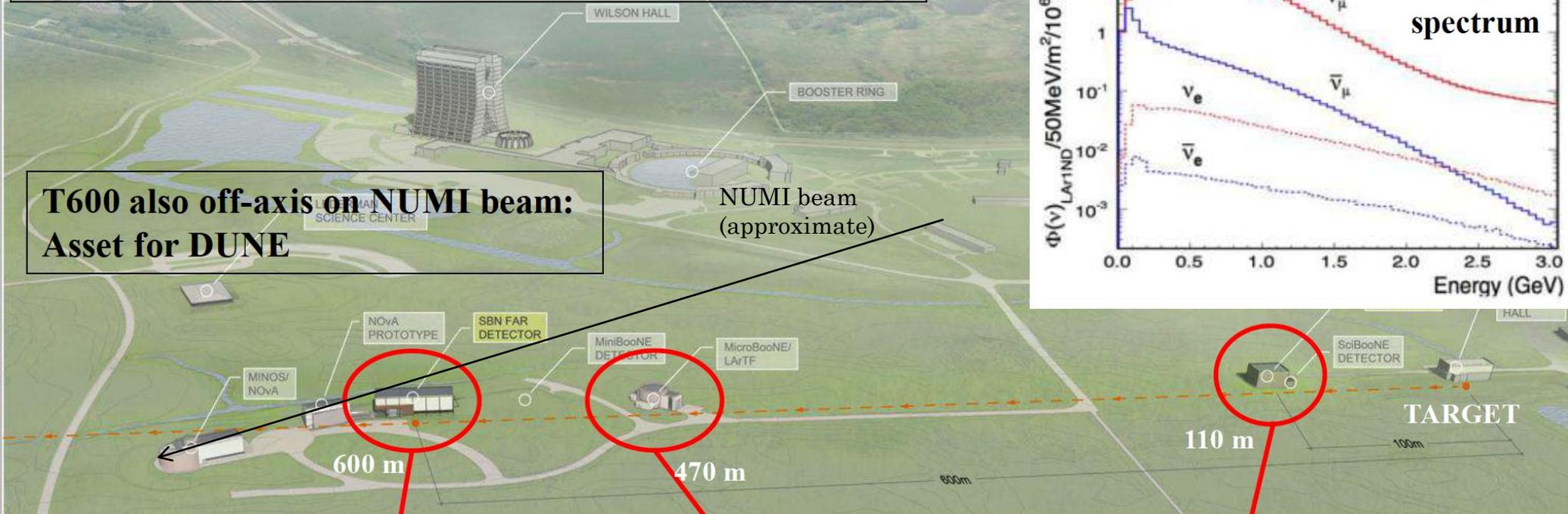
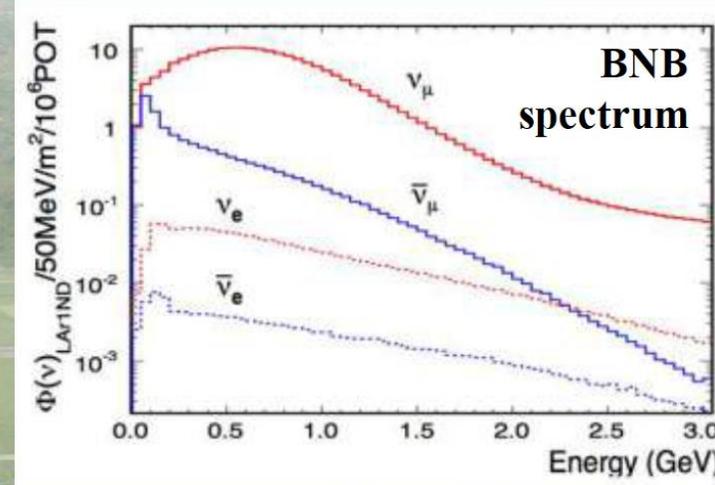
- A comparison between far/near detector is crucial for any accelerator experiment, with a better control of backgrounds and systematics

**SBN satisfies these requirements: it could have a crucial role in solving the sterile neutrino puzzle!**

# The SBN project

$$L/E_\nu \sim 600 \text{ m} / 700 \text{ MeV} \sim \mathcal{O}(1 \text{ m/MeV})$$

**T600 also off-axis on NUMI beam:  
Asset for DUNE**



**ICARUS T600**

**FAR DETECTOR:  
T600 – 476 ton**

**ICARUS**

**MicroBooNE  
89 ton**

**NEAR DETECTOR:  
SBND – 112 ton**

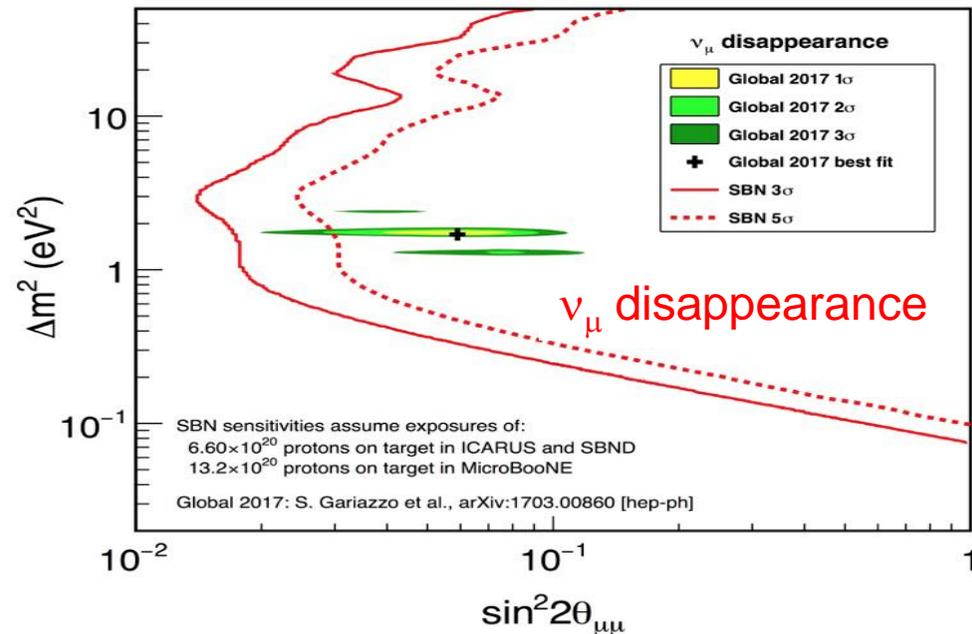
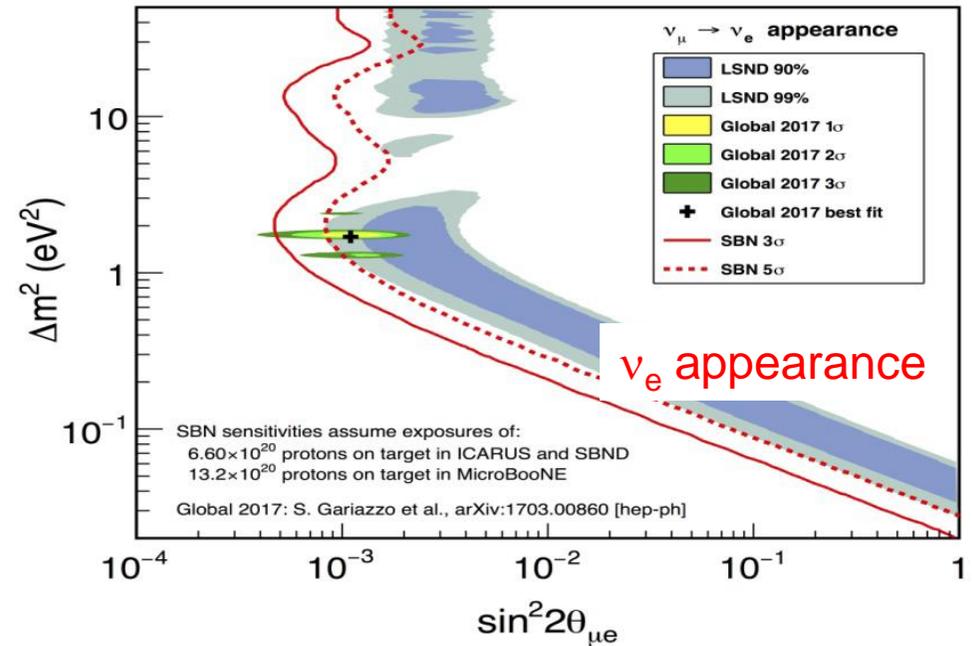


- The experiment will exploit 3 LAr-TPCs exposed to the FNAL Booster Neutrino Beam, (only  $\sim 0.5\%$   $\nu_e$  contamination) at different distances from target:  
SBND, MicroBooNE and ICARUS at 110, 470, and 600 meters respectively
- The experiment is expected to clarify the sterile anomaly by precisely/independently measuring both  $\nu_e$  appearance and  $\nu_\mu$  disappearance
- Using the same detector technology for all the 3 detectors will greatly reduce the systematic errors: SBND (near detector) will provide the "initial" beam composition and spectrum
- The great  $\nu_e$  identification capability of LAr-TPC will help reduce the NC background
- During SBN operations, ICARUS will also collect  $\sim 2$  GeV neutrinos from NuMI (Neutrino Main Injector) Off-Axis beam. This will be an asset for the future long-baseline project:
  - $\nu$  interaction cross-section measurements and identification/reconstruction studies
  - In particular, a large  $\nu_e$  component with  $\sim 3$  GeV energy (in the DUNE range)

$\nu_e$  appearance sensitivity for 3 yr  
( $6.6 \times 10^{20}$  pot) at BNB

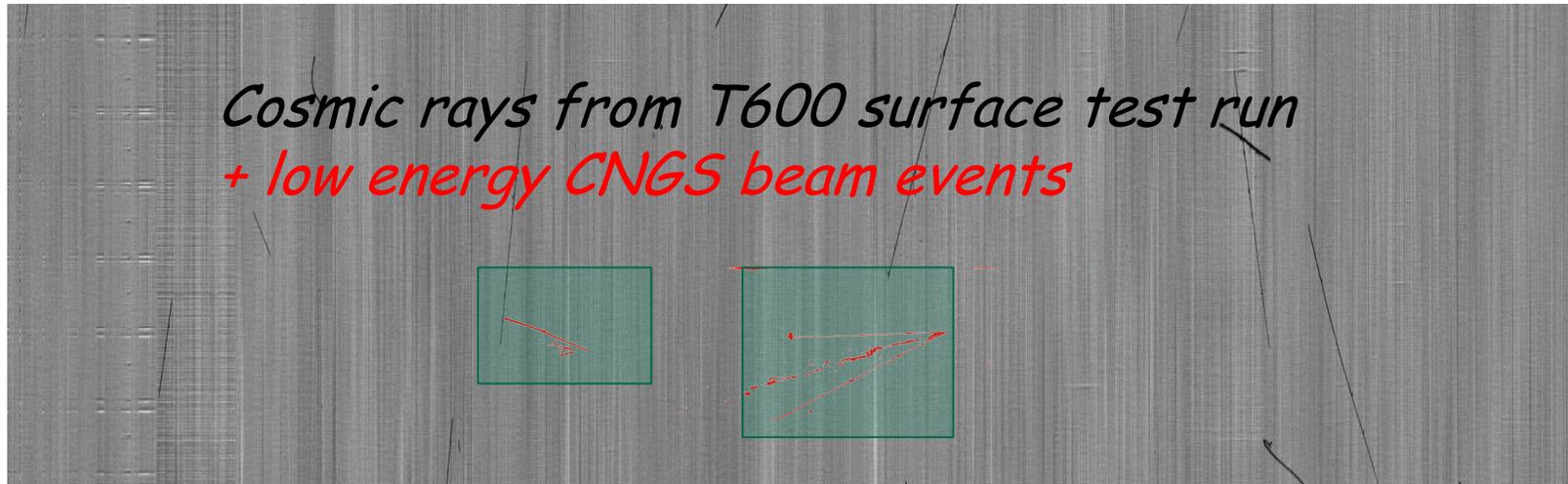
LSND 99%CL region will be  
covered at  $\sim 5\sigma$  level

Exploiting high rates and detector  
correlations,  $\nu_\mu$  disappearance  
sensitivity will be extended by **one  
order of magnitude** beyond present  
limit in 3 yr



# The new experimental challenge: a LAr-TPC on surface 13

- ICARUS will take data at shallow depth (only 3 m concrete overburden)
- ~ 11 muon tracks will hit each ICARUS module in the ~1 ms drift window and be randomly overlapped to the beam neutrino interaction



- Cosmic-related photons can mimic  $\nu_e$  interactions (via Compton or asymmetric pair production) and result in significant background
- They can be mitigated by unambiguously identifying the timing of each (cosmic and beam) ionizing event occurring during the  $O(1)$  ms TPC drift window
- This requires signals on a faster time scale than the TPC including:
  - improved LAr scintillation light detection system
  - Cosmic Ray Tagger (CRT) surrounding the detector

- To face the new experimental conditions at FNAL (shallow depth, higher beam rate) T600 underwent intensive overhauling at CERN, before shipping to US.
- Overhauling took place in the [CERN Neutrino Platform](#) framework (WA104) from 2015 to 2017.
- Several technology developments were introduced *while maintaining the already achieved performance:*
  - new cold vessels, with a purely passive insulation;
  - renovated LAr cryogenics/purification equipment;
  - improvement of the cathode planarity
  - new faster, higher-performance read-out electronics;
  - upgrade of the PMT system: higher granularity and ns time resolution

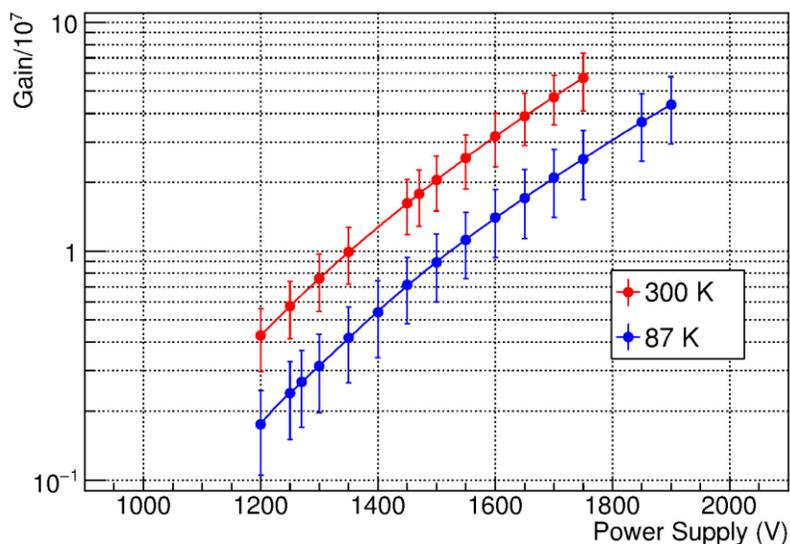


In shallow depth operation, the light collection system will allow to:

- Precisely identify the **time of occurrence ( $t_0$ )** of any ionizing event in the TPC
- Determine the event **rough topology** for selection purposes
- Generate a **trigger signal for read-out**

**ICARUS@SBN** uses 90 8" PMTs per TPC (5% coverage, 15 phe/MeV) that provides:

- Sensitivity to low energy events ( $\sim 100$  MeV)
- Good spatial resolution ( $\leq 50$  cm)
- $\approx$  ns timing resolution
- Possible cosmics identification by PMT space/time pattern



The gain reduction in LAr w.r.t. room temperature (up to a factor 10) will be compensated by a  $\sim 100$  V increase in power supply voltage

ICARUS electronics at LNGS was based on:

- "warm" low-noise front-end amplifier
- Multiplexed 10-bit ADC
- Digital VME module for local storage, data compression, trigger information

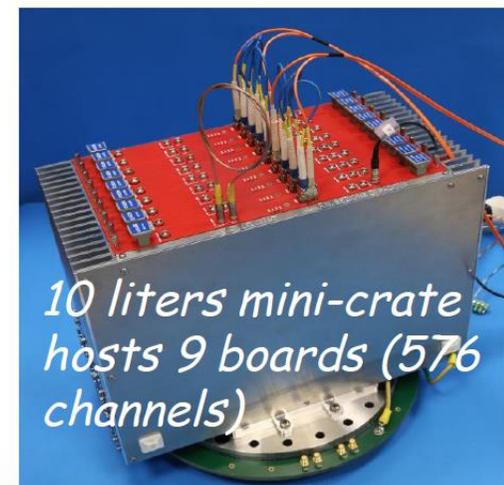
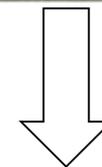
However, in view of the SBN program, some components were modernized and improved:

- Serial 12-bit ADC, fully synchronous in the whole detector  
--> ~20% improvement in muon momentum resolution via MCS
- Serial bus architecture increases transmission rate to Gbit/s
- More compact layout: both analog+digital electronics hosted on a single flange

Volume 60 times smaller

*New electronics extensively tested on a 50-liter TPC@CERN*

*JINST 13 (2018) P12007*



- T600 installed inside warm vessel in August 2018
- Installation of TPC/PMT feedthrough flanges and connectivity tests, completed by February 2019
- Leak tightness tests completed
- Top cold shields and top CRT support installed

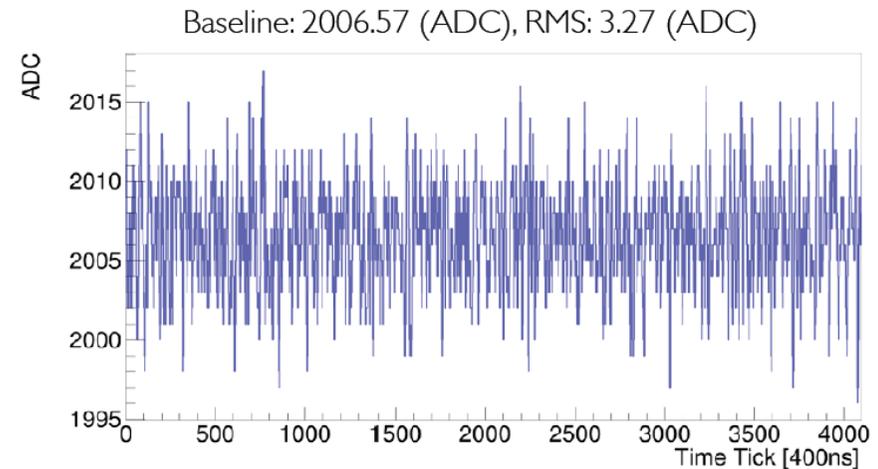


- PMT electronics and trigger being tested at cryogenic temperatures at CERN
- Installation of proximity cryogenics started in February
- Side CRT installation also ongoing
- Director's Review in December 2018 recognized the great progress of SBN

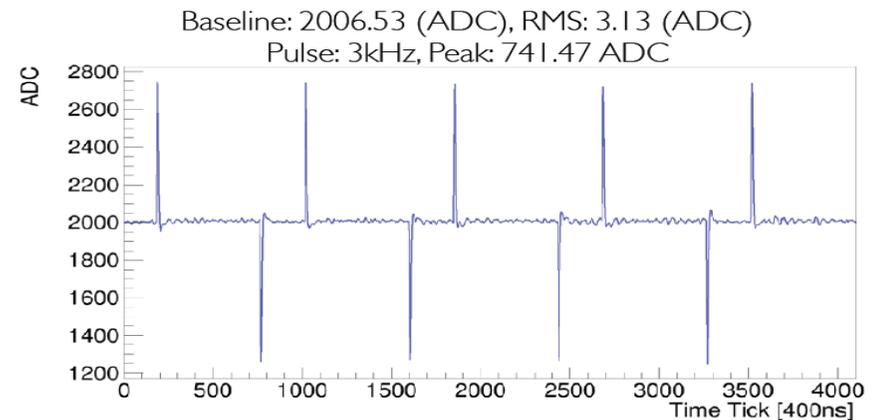
- A test of the full readout chain, from wires to DAQ, was performed in December on one mini-crate (576 channels + optical links)
- Allowed to check readout and set baseline for future noise monitoring
- Noise measured on random triggers and test pulses
- Grounding conditions different from ICARUS data-taking conditions and far from optimal
- Noise RMS  $\sim 1700 e^-$ , not too far from  $\sim 1200 e^-$  measured in CERN 50-liter setup

*The successful readout test confirms the good performance of the full TPC electronics!*

## Noise Waveform



## Injected Pulses



- Surrounds the cryostat with two layers of plastic scintillators: 1100 m<sup>2</sup>
- Tags incident cosmic or beam-induced muons with high efficiency (95%) giving spatial and timing coordinates of the track entry point
- Reconstructed CRT hits are matched to activity in the LAr volume
- Few ns time resolution allows measuring direction of incoming/outgoing particle propagation via time of flight

## TOP

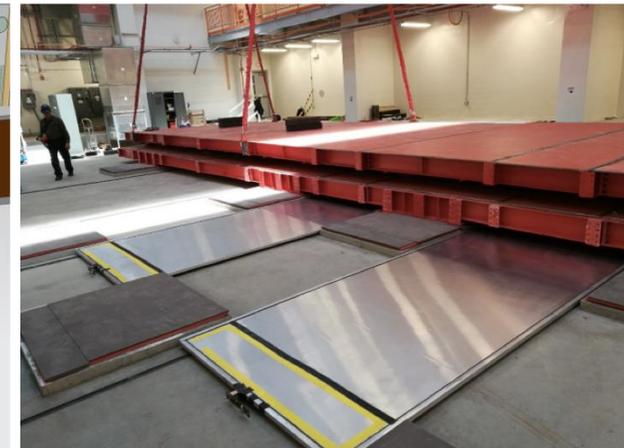
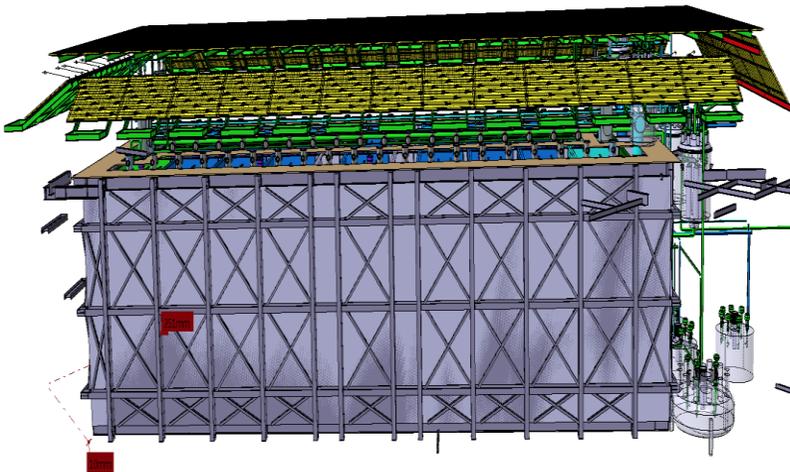
*~ 400 m<sup>2</sup>: roof+angled parts  
Will catch ~80% cosmic ms  
SiPM readout*

## SIDES

*~ 500 m<sup>2</sup> on four sides  
SiPM readout*

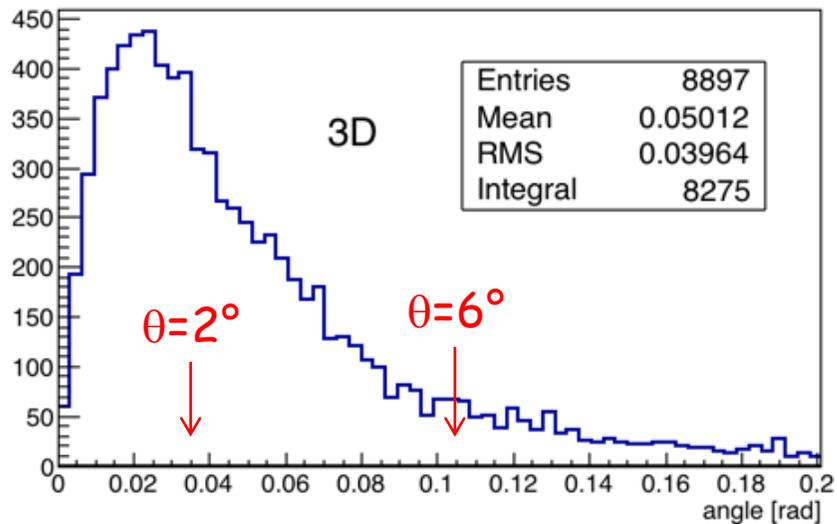
## BOTTOM

*~ 200 m<sup>2</sup>, already installed  
2 parallel layers  
PMT readout*

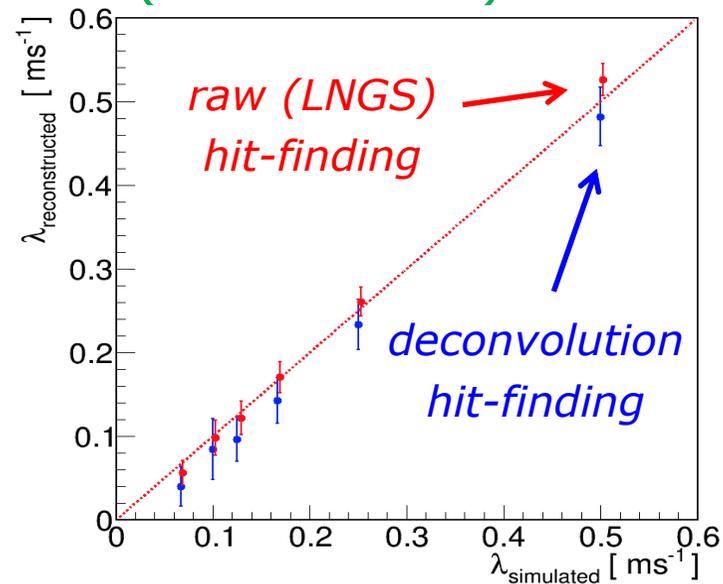


- A detailed understanding of detector-related systematics and their correlation across near/far detectors **will be crucial** to SBN physics
- **Common reconstruction tools** and oscillation analysis are therefore fundamental
- ICARUS joined the **LArSoft** framework: mutual sharing of algorithms and tools and cross-check between different reconstruction approaches

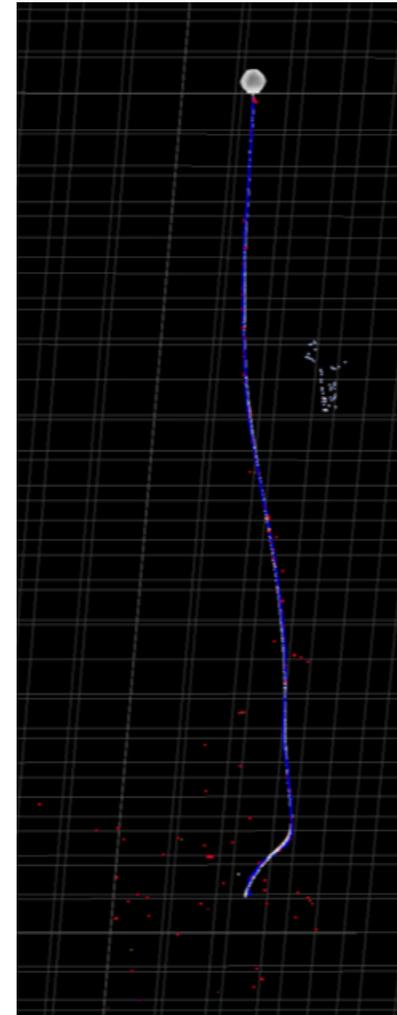
angle between sim/reco direction for EM showers



electron lifetime (reco vs. sim)



3D reco of stopping muon





*leaving from CERN, June 12<sup>th</sup> 2017*



*Antwerp: unloading from barge from Basel and loading into ship to Burns Harbor (Indiana)*



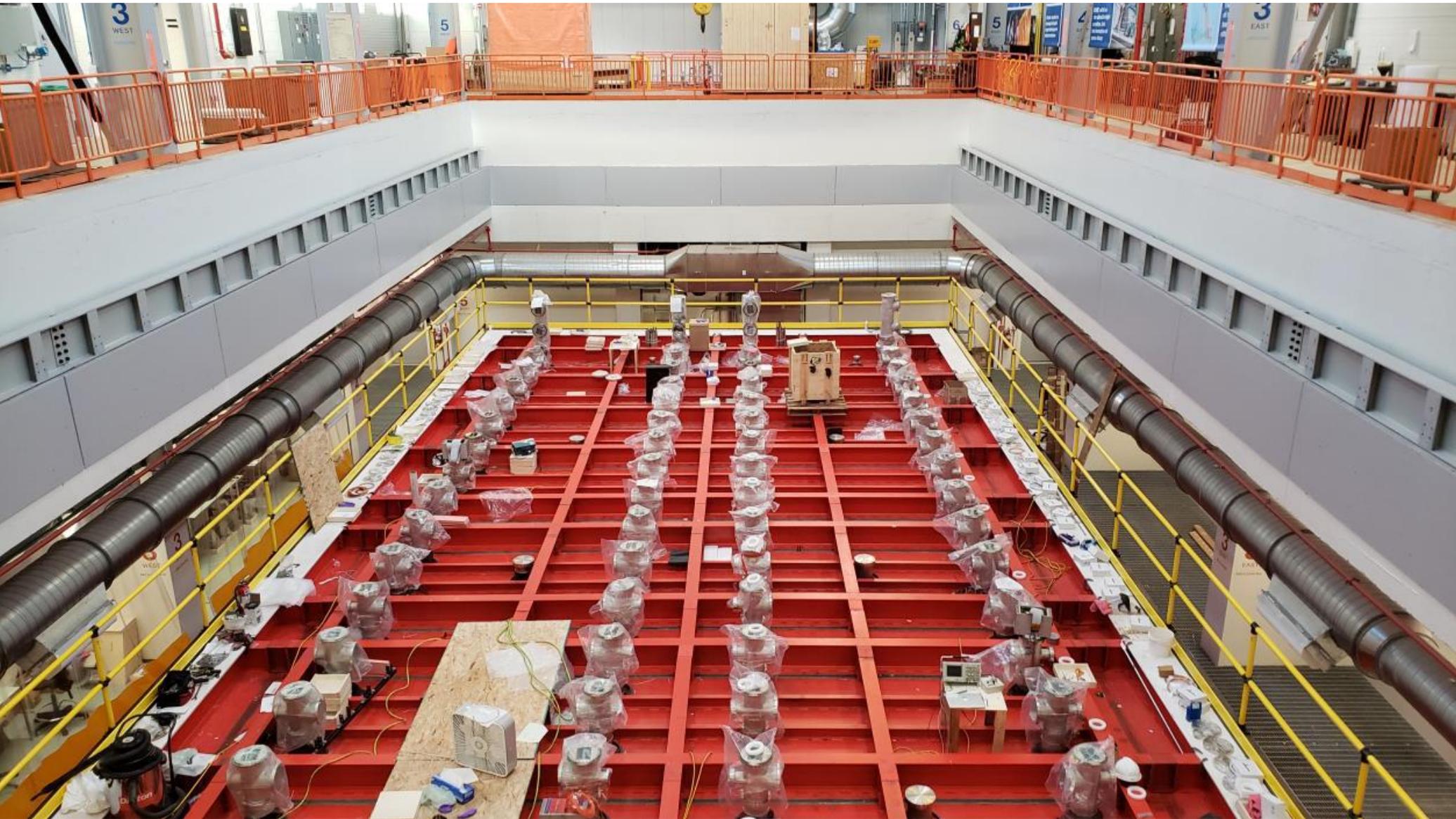
*arriving at SBN Far site building at FermiLab, July 26<sup>th</sup> 2017*



*SBN far site building at FermiLab*





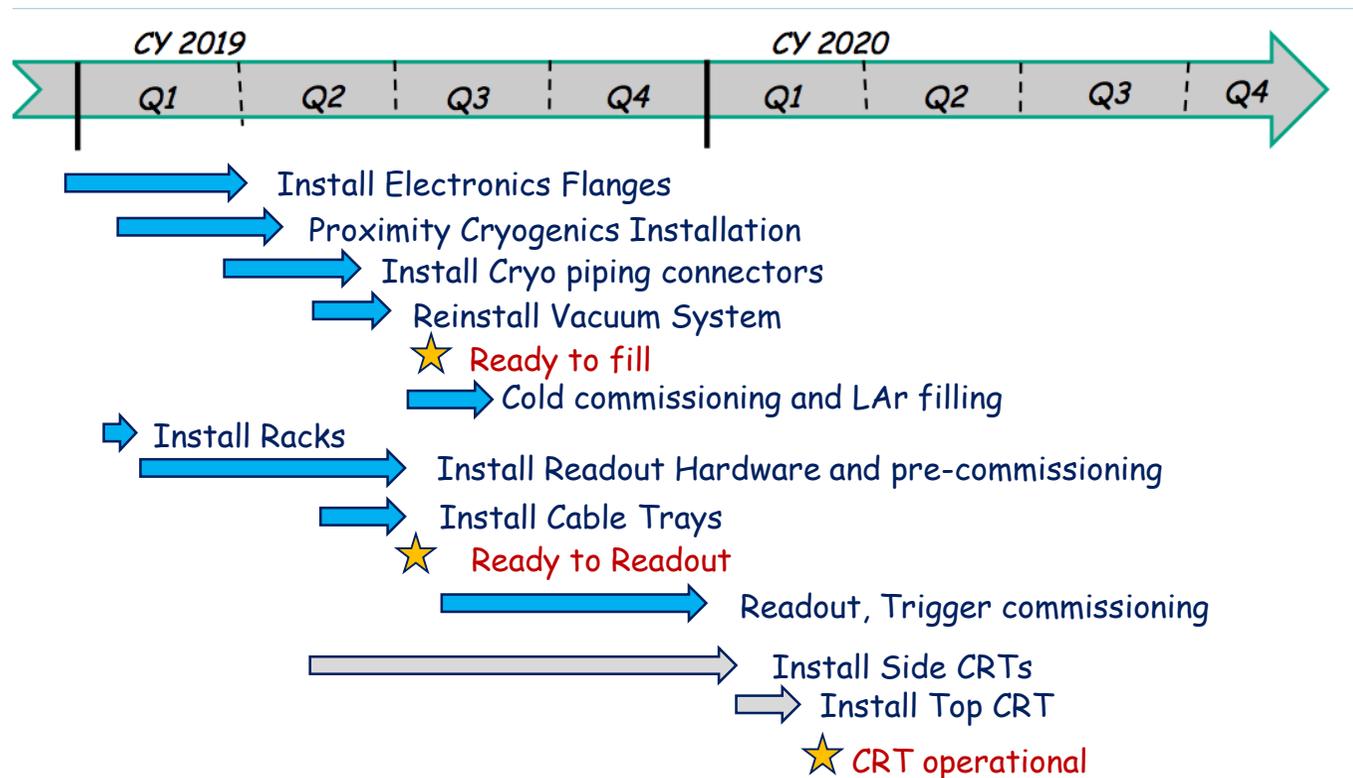




Readout electronics

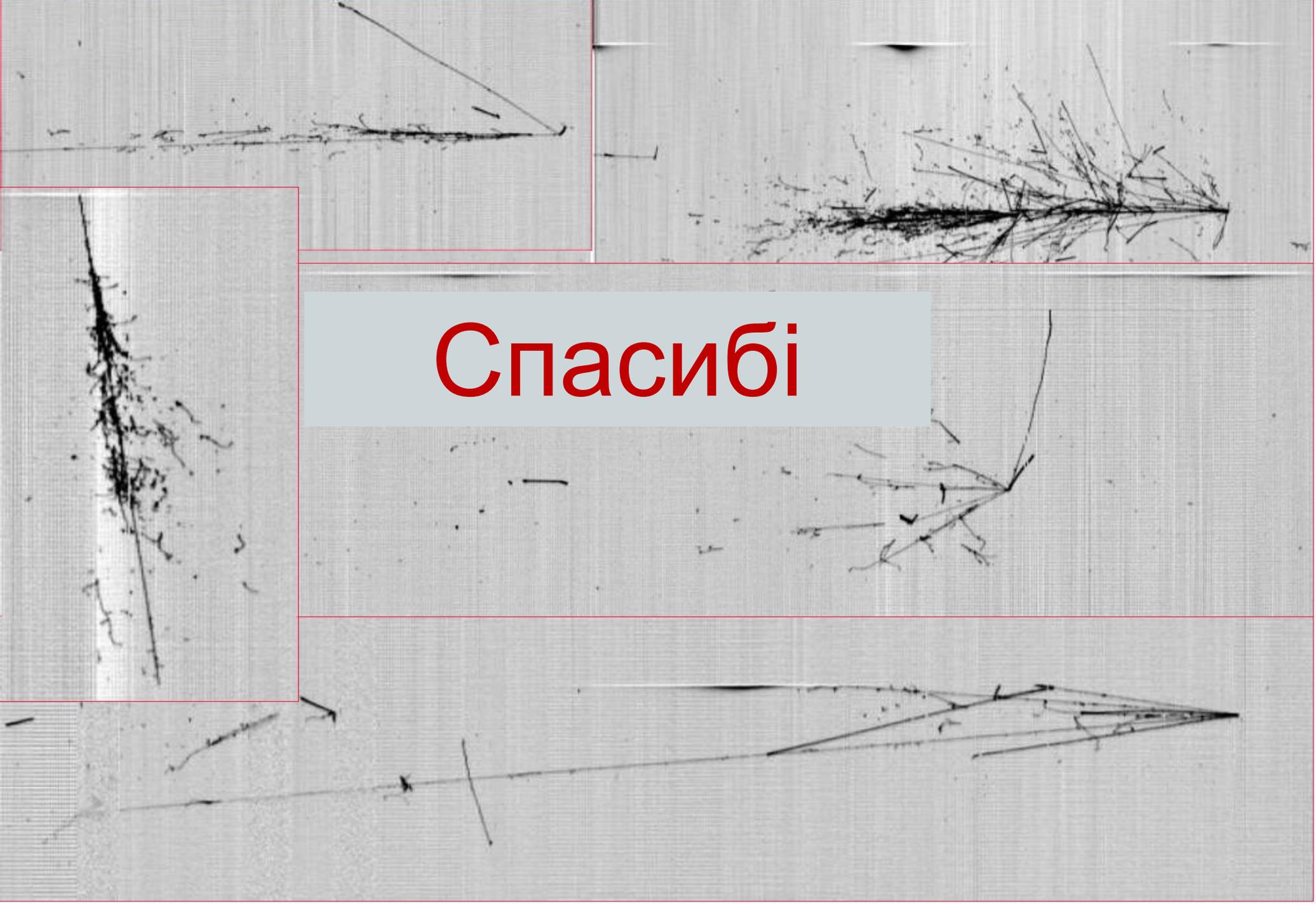
Power supply

- TPC/trigger electronics installation to be completed and tested by spring 2019
- PMT electronics installation also to be completed during the spring
- ICARUS expected to be ready to fill at the end of July 2019
- After cryogenics commissioning, cooldown and filling, ICARUS T600 should be operational during the last quarter of 2019
- Commissioning of CRT, DAQ, trigger and slow controls will follow
- Data-taking for physics is expected by the end of this year



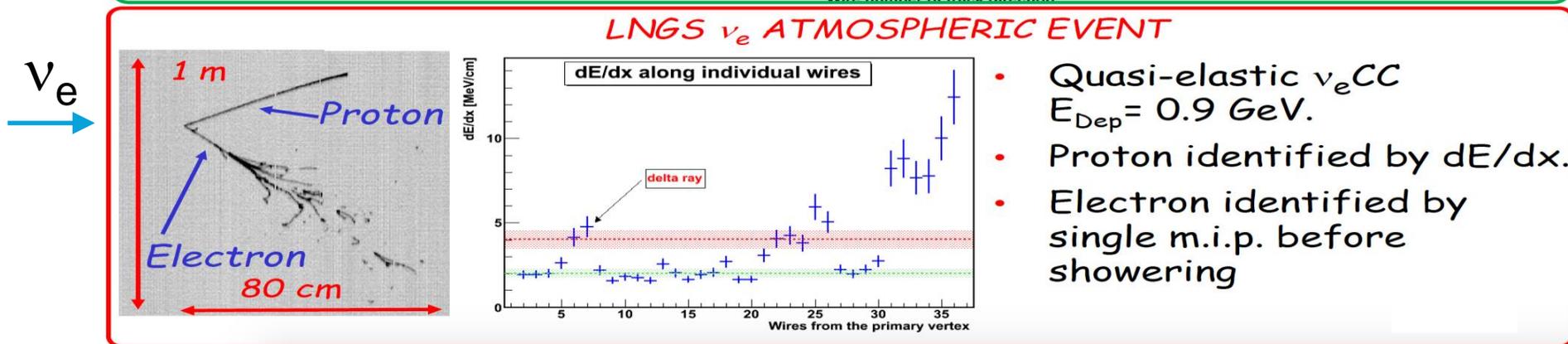
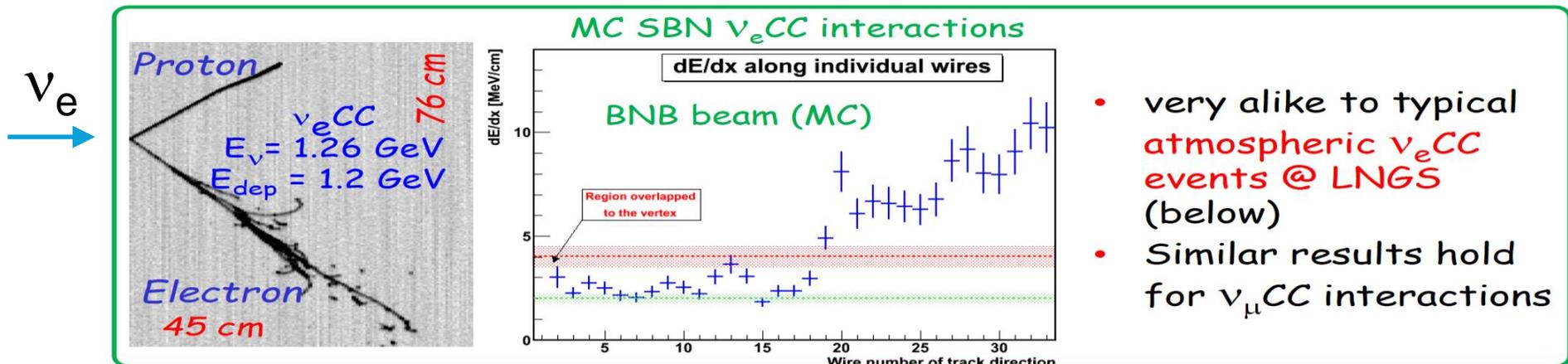
- The ICARUS-T600 successful 3-year run at LNGS proved that LAr-TPC technology is mature and ready for large-scale neutrino physics experiments
- ICARUS searched for LSND-like anomaly via  $\nu_e$  appearance in the CNGS beam. The negative result constrained significantly the allowed parameter region
- The SBN project at FNAL is expected to clarify the sterile neutrino puzzle, by looking at both appearance and disappearance channels with three LAr-TPCs
- After an extensive refurbishing, ICARUS is being installed as the SBN far detector at FNAL. Data taking expected in 2019, near detector in 2021
- ICARUS will see first neutrinos by the end of this year !

Спасибі

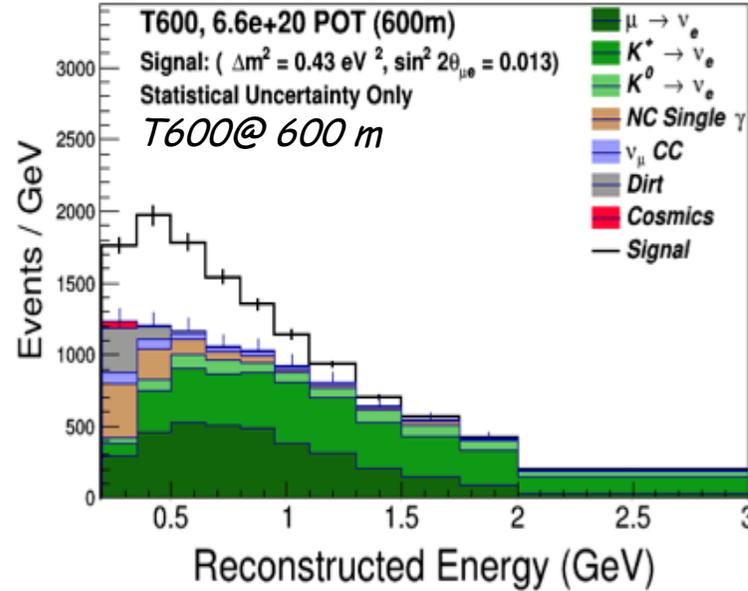
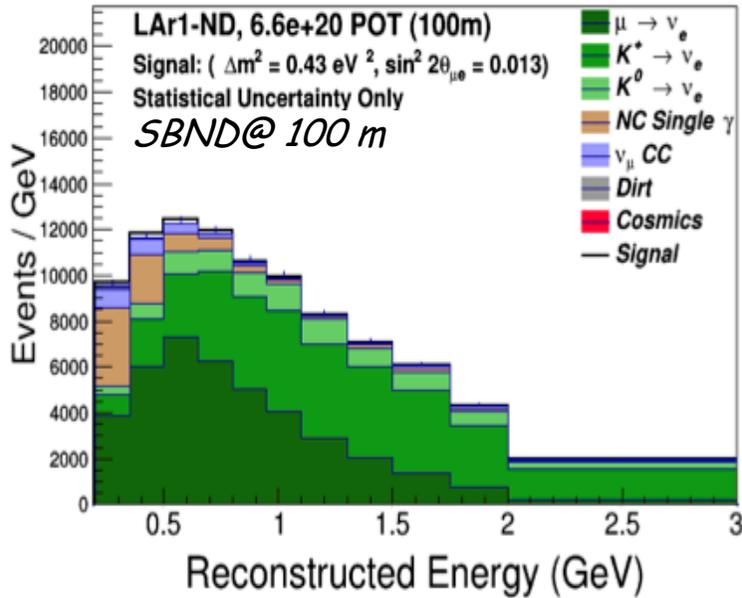


Backup slides

- ICARUS at LNGS was also exposed to atmospheric neutrinos (exposure  $\sim 0.74$  kt year) first observation of atmospheric neutrinos with a LAr-TPC
- 14 events found (8  $\nu_e$  CC + 6  $\nu_\mu$  CC) vs. 18 expected (taking into account: triggering, filtering and scanning efficiencies)
- Very good benchmark for the forthcoming SBN experiment: similar energy/features. Useful to develop filtering and reconstruction tools



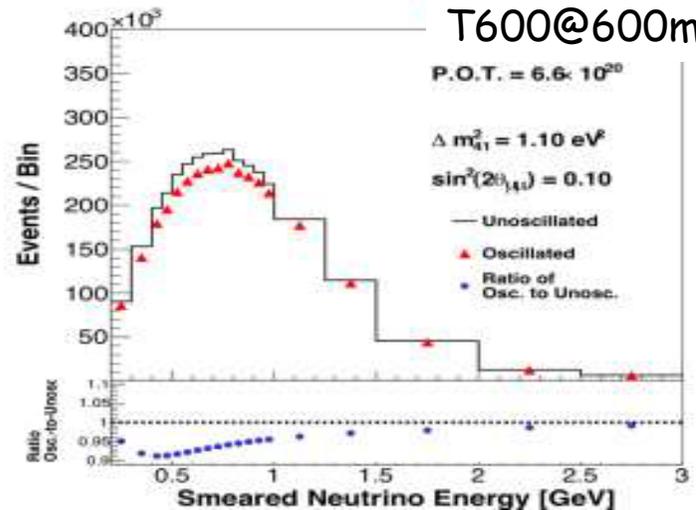
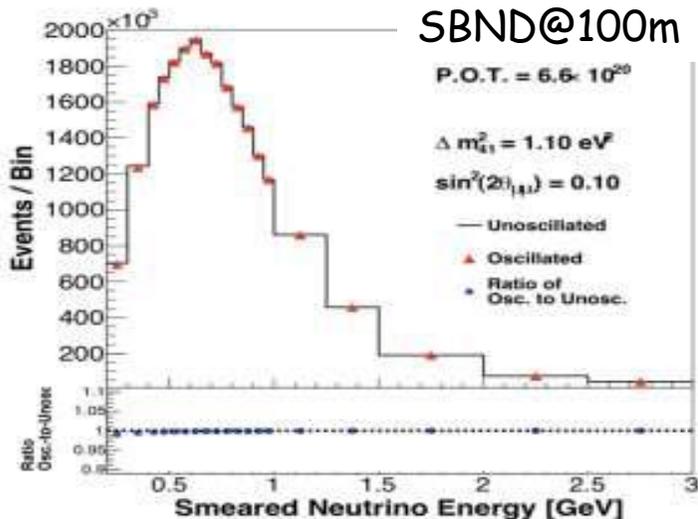
*$\nu_e$  spectra (oscillation signal + backgrounds) for 3 years ( $6.6 \cdot 10^{20}$  pot)*



Example for  
 $\sin^2 2\theta = 0.013$   
 $\Delta m^2 = 0.43 \text{ eV}^2$

In absence of  
 oscillations,  
 spectra should  
 be ~ identical

*$\nu_\mu$  spectra (oscillation modulation) for 3 years ( $6.6 \cdot 10^{20}$  pot)*



Example for  
 $\sin^2 2\theta = 0.01$   
 $\Delta m^2 = 1.10 \text{ eV}^2$

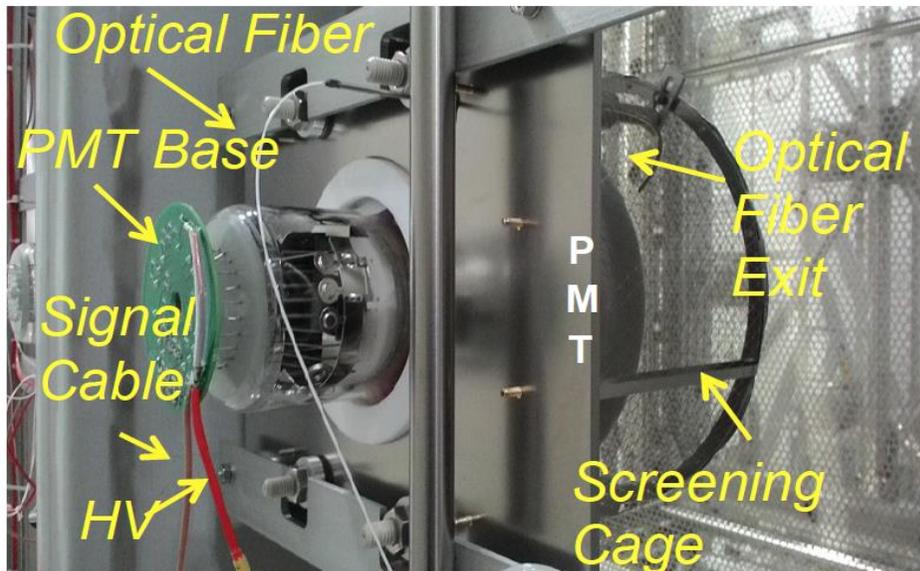
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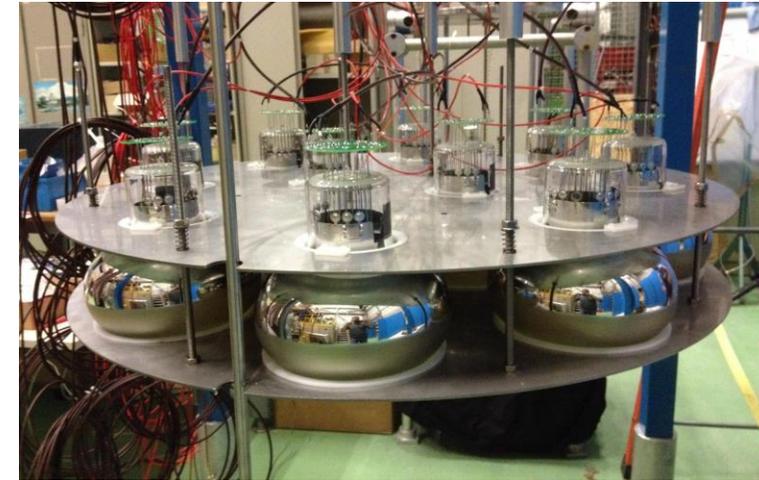
Timing/gain equalization will be performed with laser pulses

$\lambda = 405$  nm

FWHM  $< 100$  ps

peak power  $\sim 400$  mW

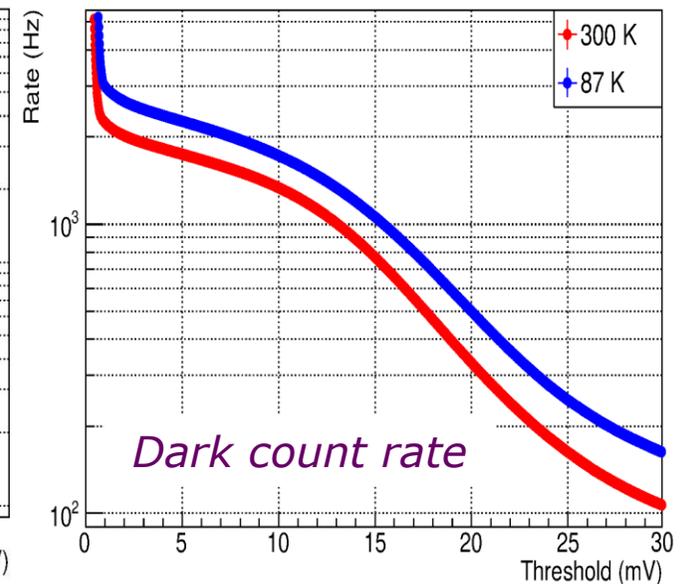
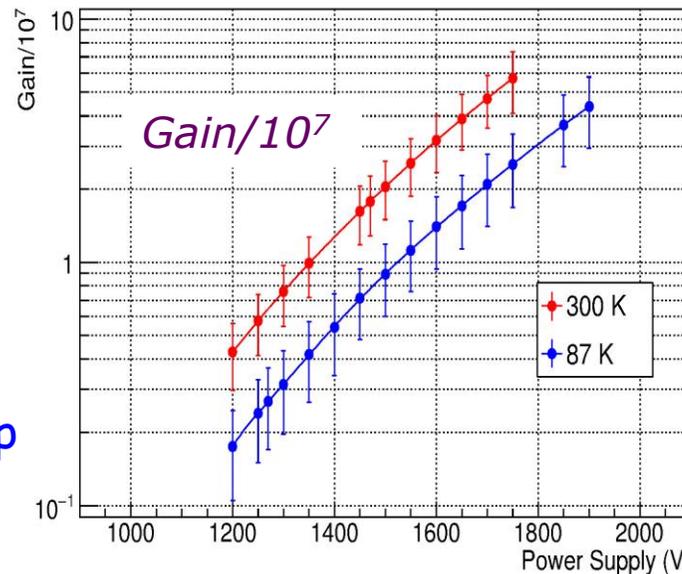
- All PMTs tested at room temperature in a dedicated dark room at CERN
- A subset of 60 PMTs tested immersed in LAr to compare the PMT performance in cryogenic environment to room temperature
- All PMTs illuminated with laser light pulses



PMTs were characterized individually at 300K and 87K:

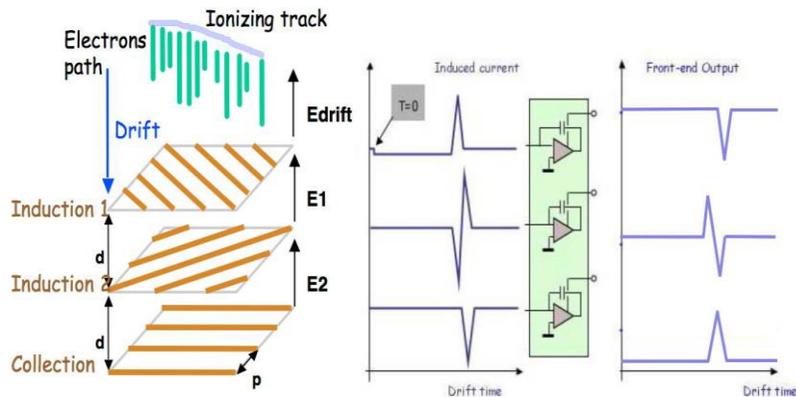
- Gain
- Dark count rate
- Peak/valley ratio
- Uniformity of photocathode response

The gain reduction in LAr w.r.t. room temperature (up to a factor 10) will be compensated by a  $\sim 100$  V increase in power supply voltage



The analog front-end shaping was also modified:

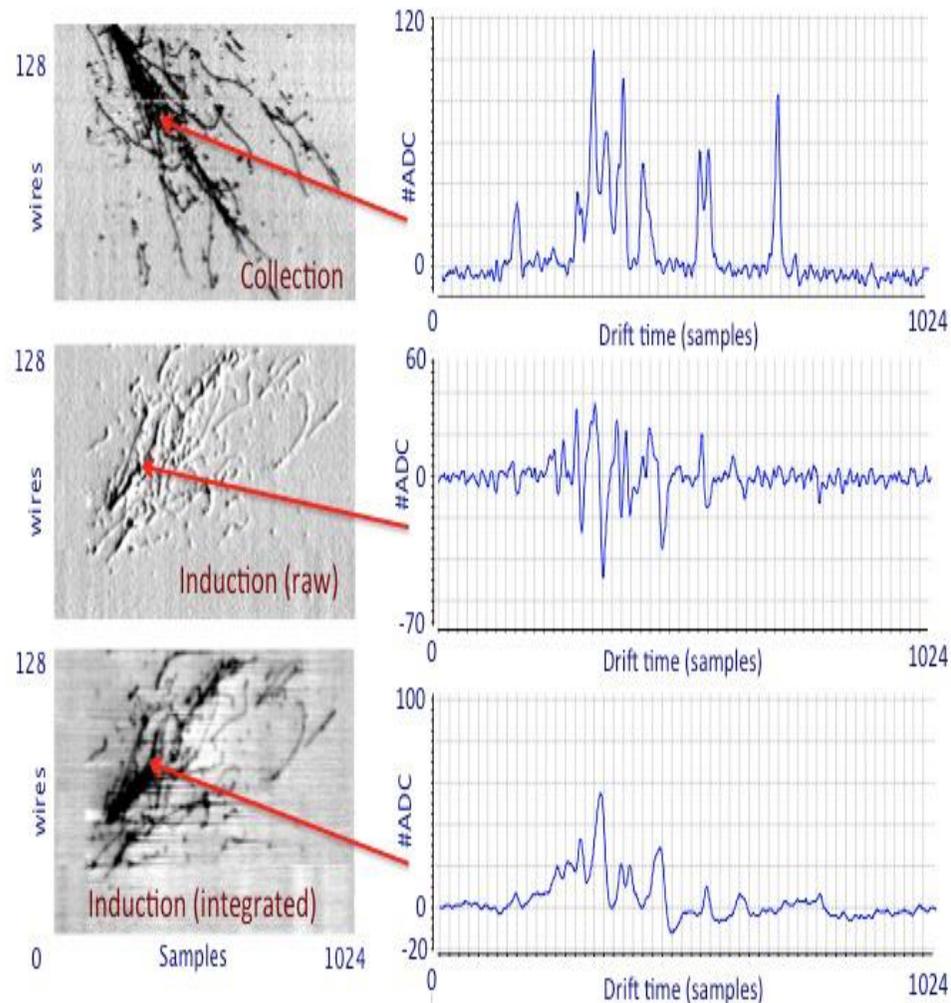
- Lower noise  $\sim 1200 e^-$  equivalent ( $\sim 20\%$  S/N improvement w.r.t. LNGS electronics)
- Shorter shaping time ( $\sim 1.5 \mu s$  for all planes) matching electron transit time between planes
- Drastic reduction of undershoot after large signals: better description of crowded vertex region



In particular, Induction 2 signal keeps bipolar shape (unlike in old front-end)

- Possible off-line integration with suitable LF filtering
- Allows calorimetric measurement in this plane too (with  $\sim 2$  worse resolution than Collection)
- May improve  $\nu_e$  identification efficiency by  $\sim 20\%$

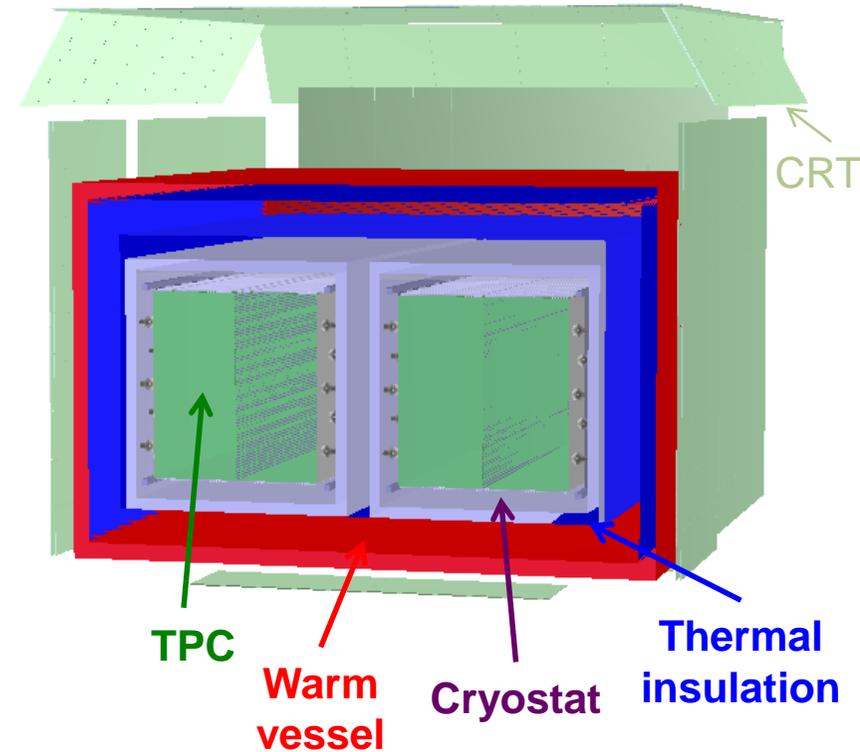
## Tests on 50-liter TPC at CERN:



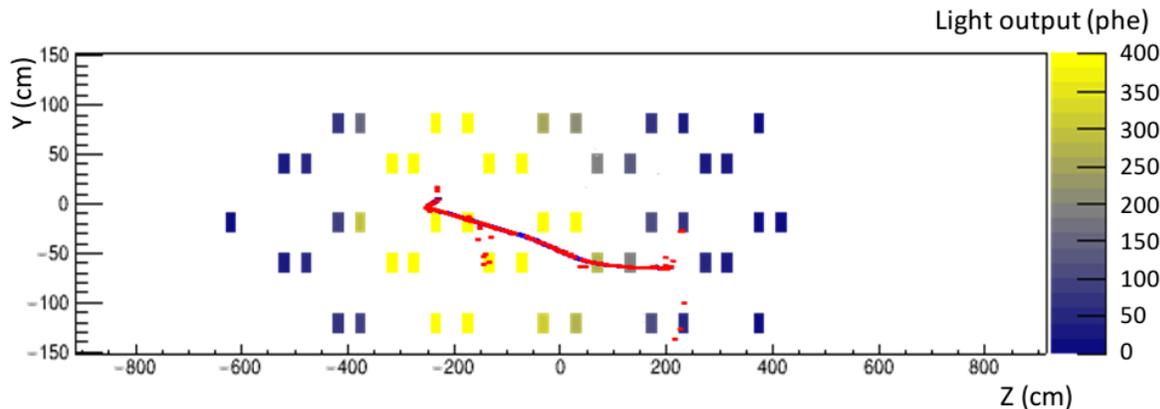
# ICARUS simulation and reconstruction

- LNGS experience will be fundamental for SBN analysis
- However, different experimental conditions:
  - Much larger event statistics
  - Overlap of cosmics on  $\nu$  events
  - Rich information from light and CRT
- Full simulation performed with realistic geometry and signals from all subdetectors (TPC, PMT, CRT)
- Signal/noise modelization uses information from LNGS and small prototypes at CERN

## Scheme of detector geometry



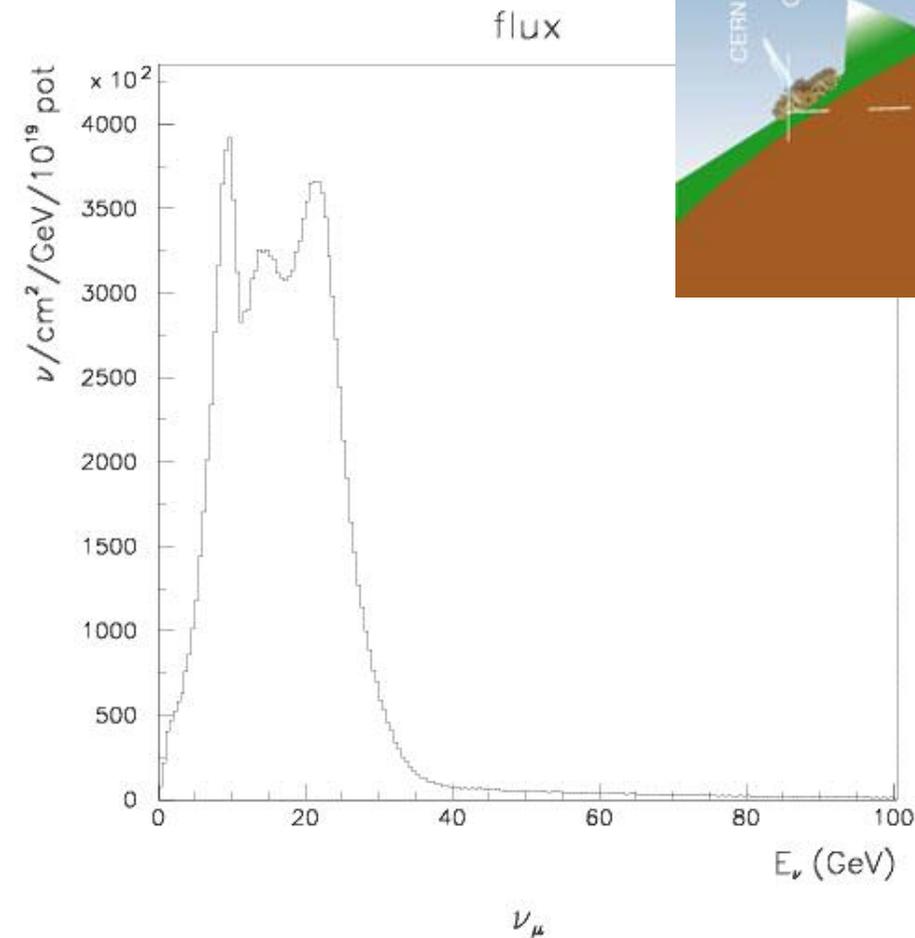
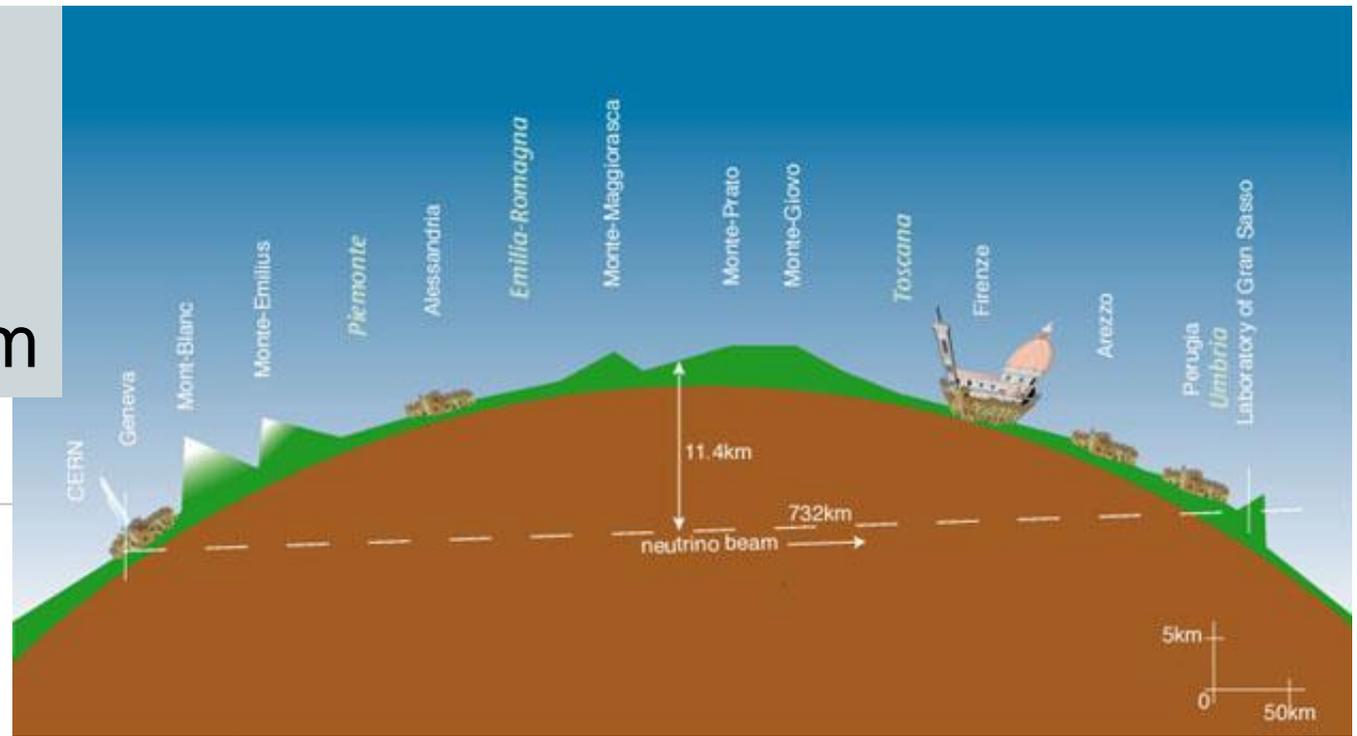
## PMT vs. TPC information



- $\nu_{\mu}$  CC interaction with  $\sim 1.4$  GeV deposited energy
- 49 PMTs have visible signal (over 10 phe)
- Fired PMTs extend over  $\sim 8$  m

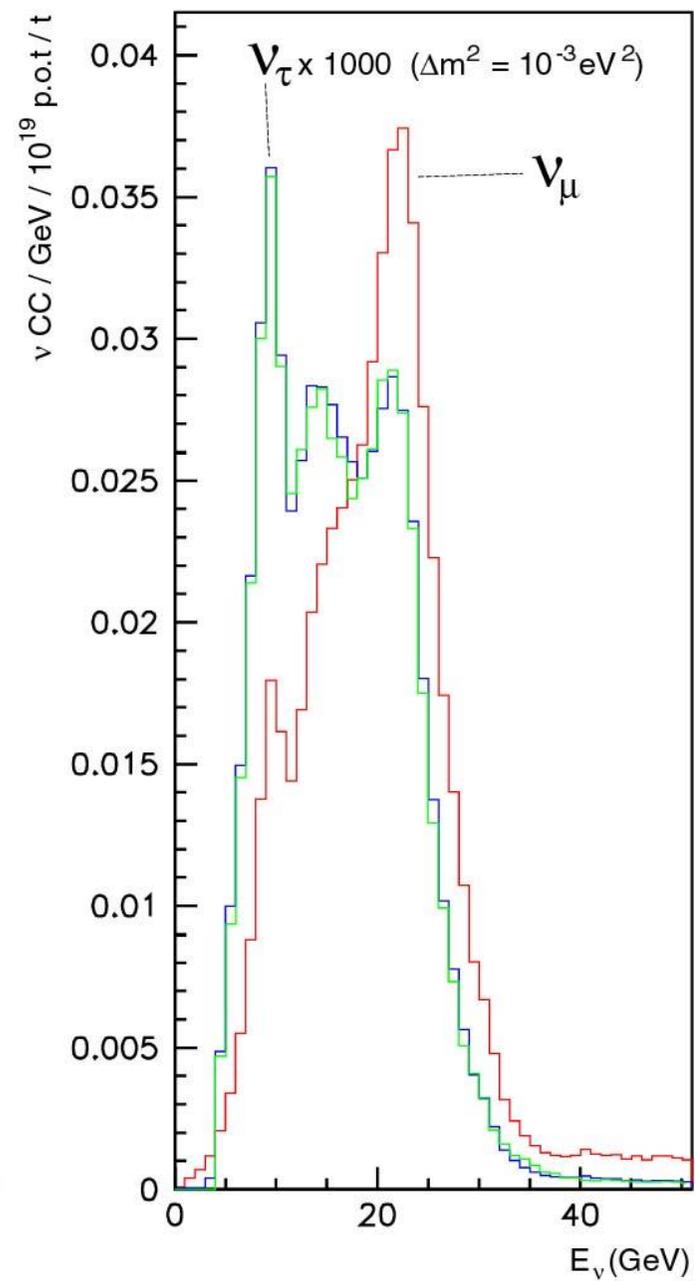
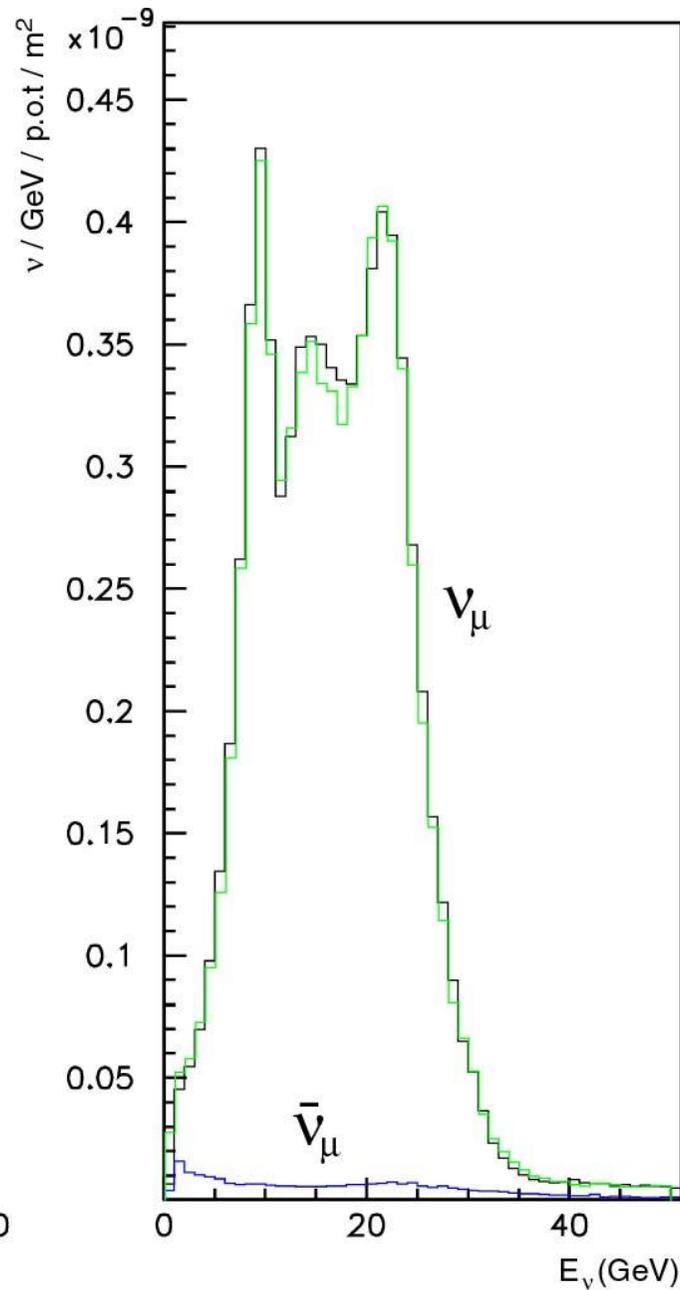
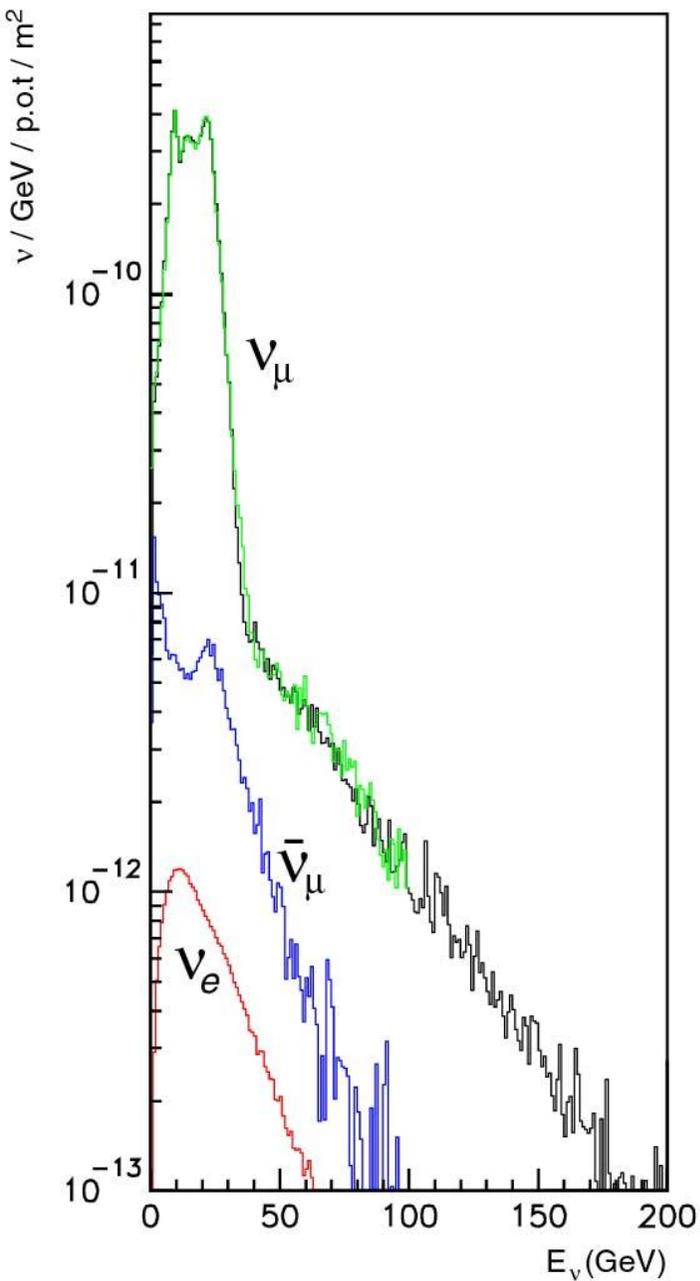
# CNGS

## CERN Neutrino to Gran Sasso beam

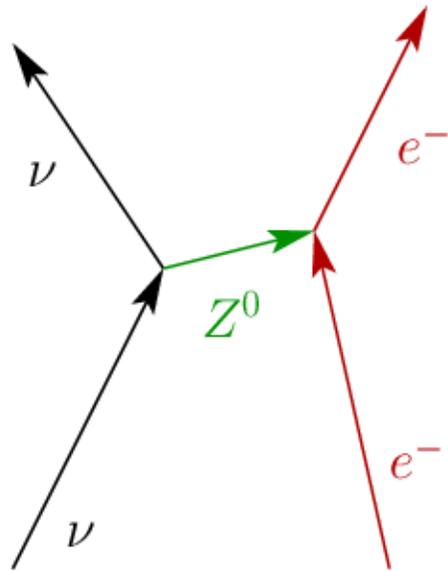


$$L/E = 730/20 \sim 36 \text{ km/GeV}$$

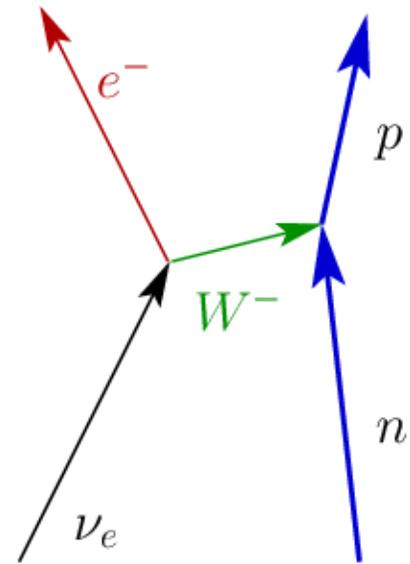
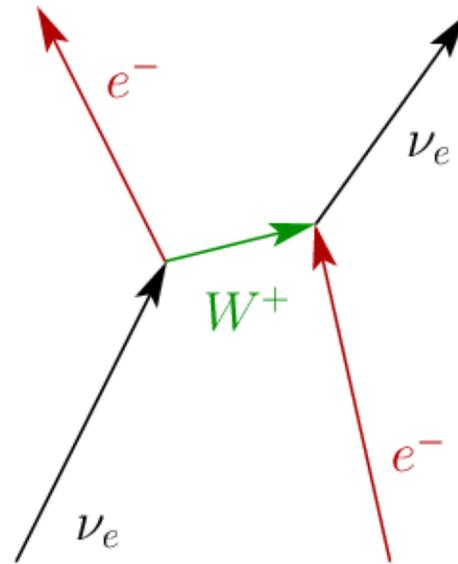
# CNGS beam spectrum



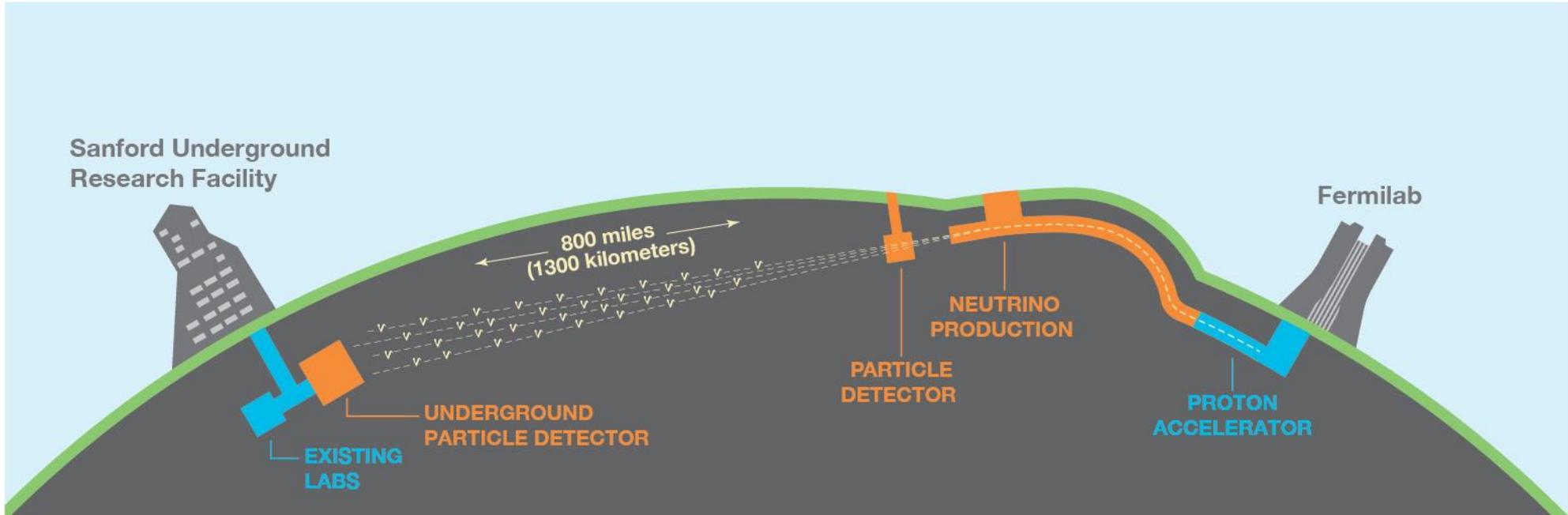
# Neutrino interactions



Neutral Current



Charged Current



**Origin of Matter** Could neutrinos be the reason that the universe is made of matter rather than antimatter? By exploring the phenomenon of neutrino oscillations, DUNE seeks to revolutionize our understanding of neutrinos and their role in the universe.



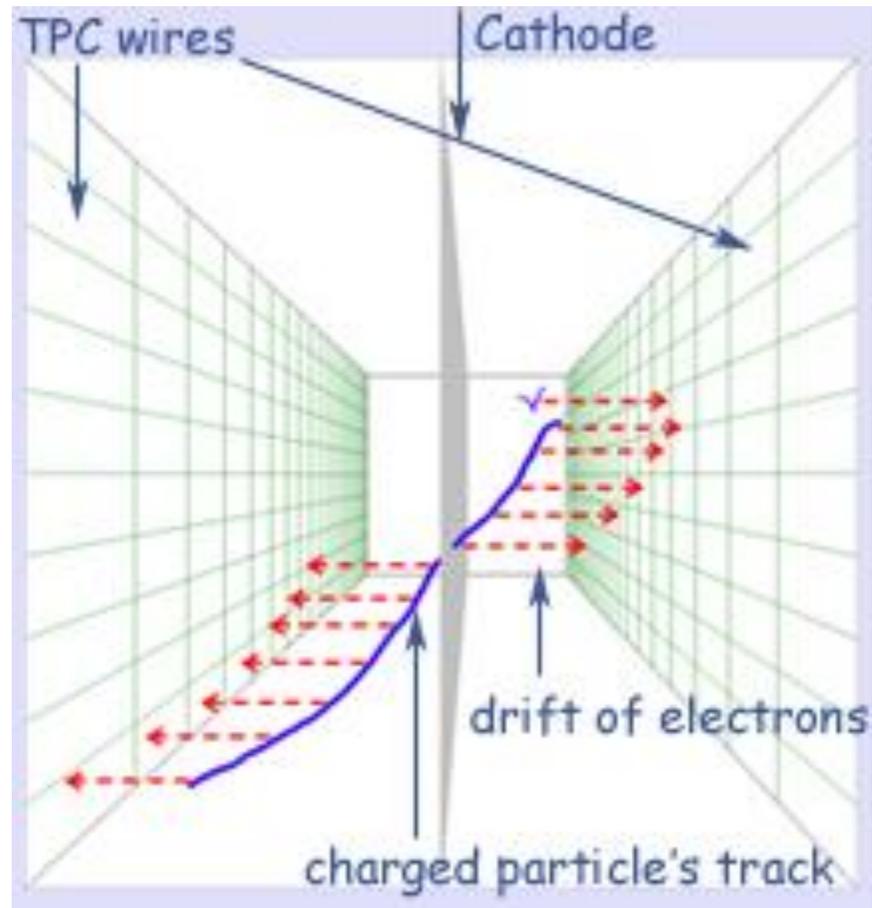
**Unification of Forces** With the world's largest cryogenic particle detector located deep underground, DUNE can search for signs of proton decay. This could reveal a relation between the stability of matter and the Grand Unification of forces, moving us closer to realizing Einstein's dream



**Black Hole Formation** DUNE's observation of thousands of neutrinos from a core-collapse supernova in the Milky Way would allow us to peer inside a newly-formed neutron star and potentially witness the birth of a black hole.

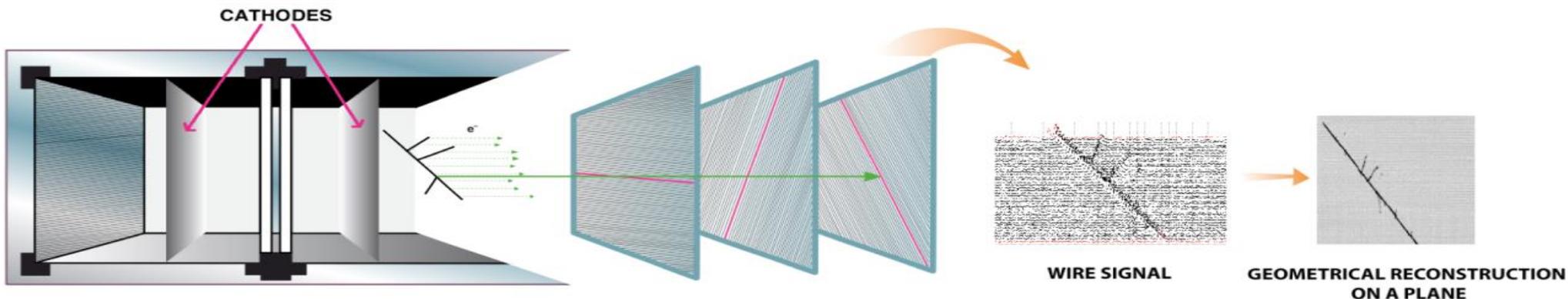
# LAr TPC principle of operation

TPC = Time Projection Chamber



# ICARUS-T600: LAr TPC working principle

- 3 wire planes with  $0^\circ$ ,  $+60^\circ$ ,  $-60^\circ$  orientation w.r.t. horizontal; both pitch between wires and distance between planes are 3 mm
- Wire biasing guarantees (almost complete) plane transparency, and non-destructive readout in both Induction wire planes
- Last plane collects charge, allowing measurement of deposited energy
- $\sim 54000$  total readout wires, with 400 ns sampling time



FRONT VIEW OF THE DETECTOR

WIRE PLANES  
ANODE

WIRE SIGNAL

GEOMETRICAL RECONSTRUCTION  
ON A PLANE

*Typical signal in  
collection (last) plan*

