

# Reconstructed potentials of the 2-field model of dark energy with canonical and non-canonical kinetic terms

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We generalize quintom to include the tachyonic kinetic term along with the classical one. For the spatially flat, homogeneous and isotropic Universe with Friedmann-Robertson-Walker metric of 4-space we discuss in detail the reconstruction of the scalar fields potential  $U(\phi, \xi)$ . Such a reconstruction cannot be done unambiguously, so we consider 3 simplest forms of  $U(\phi, \xi)$ : the product of  $\Phi(\phi)$  and  $\Xi(\xi)$ , the sum of  $\Phi(\phi)$  and  $\Xi(\xi)$  and this sum to the  $\kappa$ th power. We present the maps of reconstructed potentials for the Chevallier-Polarski-Linder parametrization of the equation of state parameter and the set of cosmological parameters obtained from the latest Planck data.

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